The Matérn Model: A Journey through Statistics, Numerical Analysis and Machine Learning

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Abstract. The Matérn model has been a cornerstone of spatial statistics for more than half a century. More recently, the Matérn model has been central to disciplines as diverse as numerical analysis, approximation theory, computational statistics, machine learning, and probability theory. In this article we take a Matérn-based journey across these disciplines. First, we reflect on the importance of the Matérn model for estimation and prediction in spatial statistics, establishing also connections to other disciplines in which the Matérn model has been influential. Then, we position the Matérn model within the literature on big data and scalable computation: the SPDE approach, the Vecchia likelihood approximation, and recent applications in Bayesian computation are all discussed. Finally, we review recent devlopments, including flexible alternatives to the Matérn model, whose performance we compare in terms of estimation, prediction, screening effect, computation, and Sobolev regularity properties.

Keywords: Approximation Theory, Compact Support, Covariance, Kernel, Kriging, Machine Learning, Maximum Likelihood, Reproducing Kernel Hilbert Spaces, Spatial Statistics, Sobolev Spaces.

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1. INTRODUCTION

This paper serves two purposes: On the one hand, we 1 provide a panoramic view, across several disciplines, of 2 the Matérn model. On the other hand, the paper provides 3 constructive criticisms about the role of the Matérn model 4 in several disciplines, while discussing alternative or more 5 general models and their relevance to many aspects of sta-6 tistical modeling, estimation, prediction, computational 7 statistics, numerical analysis, and machine learning. 8

A historical account of the Matérn model is provided
by Guttorp and Gneiting [70]. The Matérn model – also

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called the Matérn *covariance function*, or the Matérn *ker-nel*, depending on context – is commonly attributed to Matérn [108], but can be found under alternative names in different branches of the scientific literature. The use of the Matérn model is widespread, and it is impossible to cover all its diverse applications here; our review focuses on a selection of applications that are of especial interest and significance. Specifically, we aim to cover

- estimation and prediction using the Matérn model in statistics, with emphasis on maximum likelihood estimation, Kriging prediction, and the associated screening effect;
- 2. applications of the Matérn model in
 - a) computational statistics, including the stochastic differential equation (SDE) and stochastic partial differential equation (SPDE) approaches, likelihood approximation, inference of partial differential equations (PDEs) and Stein's method;

 b) statistical modeling, including non-standard scenarios, for instance when isotropy and stationarity cannot be assumed, or to model directions and curves;

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kernel-based interpolants; d) machine learning, where the Matérn model is central to the literature on Gaussian processes modelling; and

where the Matérn model is used to construct

- e) probability theory, where the Matérn model has inspired several contributions based on properties of the sample paths of associated stochastic processes, in concert with the solution of certain classes of stochastic differential equations;
- 3. comparison with recent flexible alternatives to the Matérn model, with a focus on
 - a) enhanced models with interesting features, such as compact support or polynomial decay;
 - b) asymptotic estimation accuracy, misspecified prediction, and screening effects;
 - c) the implications of using certain classes of compactly supported kernels within approximation theory, computational statistics, and machine learning.

This article is novel, in being the first to take a broad view 56 of the scientific literature through the lens of the Matérn 57 model. In particular, we do not attempt a review of covari-58 ance functions in general. Recent reviews provide a quite 59 exhaustive panorama of covariance models, from space 60 to space-time [126], to multivariate covariance functions 61 [60], and covariance-based modeling on spheres and man-62 ifolds [122]. In addition, while there are many fascinat-63 ing applications of the Matérn model across the scientific 64 landscape, we cannot hope to do justice to them all. Our 65 emphasis is therefore limited to methodological and the-66 oretical issues which we hope are of relevance across a 67 wide range of disciplines in which the Matérn model is 68 used. 69

1.1 Setting and Notation 70

Throughout, bold letters refer to vectors and matrices, and the transpose operator is denoted \top . Let $d \in \mathbb{N}$ and let $Z = \{Z(\boldsymbol{x}), \, \boldsymbol{x} \in \mathbb{R}^d\}$ be a real-valued Gaussian random field, having zero mean and and *covariance function* K: $\mathbb{R}^d \times \mathbb{R}^d \to \mathbb{R}$ defined via $K(\boldsymbol{x}, \boldsymbol{y}) := \operatorname{Cov}(Z(\boldsymbol{x}), Z(\boldsymbol{y})).$ Covariance functions are symmetric and positive definite, where in this paper the term *positive definite* is understood as

(1)
$$\sum_{i=1}^{n} \sum_{j=1}^{n} c_i K(\boldsymbol{x}_i, \boldsymbol{x}_j) c_j \ge 0$$

for all $c_i \in \mathbb{R}$, all $n \in \mathbb{N}$ and all $x_i \in \mathbb{R}^d$. If the inequal-71 ity above is strict, then K will be called strictly positive 72 definite. 73

Each symmetric positive definite function $K : \mathbb{R}^d \times$ $\mathbb{R}^d \to \mathbb{R}$ defines *translate* functions $K(\boldsymbol{x}, \cdot)$ on \mathbb{R}^d , for all $x \in \mathbb{R}^d$. In addition, one can define an inner product on two translates by

(2)
$$\langle K(\boldsymbol{x},\cdot), K(\boldsymbol{y},\cdot) \rangle_{\mathcal{H}(\mathcal{K})} := K(\boldsymbol{x},\boldsymbol{y}), \ \boldsymbol{x}, \boldsymbol{y} \in \mathbb{R}^d$$

in terms of K itself. This extends to all linear combinations of translates and generates, by completion, a Hilbert space $\mathcal{H}(\mathcal{K})$ of functions on \mathbb{R}^d . This space is called the *native* space for K. Notice that the Hilbert space allows for continuous point evaluations $\delta_{\boldsymbol{x}} : f \mapsto f(\boldsymbol{x})$ via a *re*production formula

(3)
$$f(\boldsymbol{x}) = \langle f, K(\boldsymbol{x}, \cdot) \rangle_{\mathcal{H}(K)}, \ \boldsymbol{x} \in \mathbb{R}^d, \ f \in \mathcal{H}(K)$$

that follows from (2). Then $\mathcal{H}(K)$ is called a *reproducing* kernel Hilbert space (RKHS) with kernel K. In particular, the translates $K(\boldsymbol{x}, \cdot)$ lie in $\mathcal{H}(K)$, forming its completion and being the Riesz representers of delta functionals δ_x . They are central to machine learning, numerical analysis and approximation theory, since (2) allows inner products in the abstract space $\mathcal{H}(K)$ to be explicitly computable using the kernel - the so-called kernel trick. See Section 6.1 and [166] for more detail. For a positive definite and stationary kernel K, its Fourier transform K can be used to recast the inner product (2) on the Hilbert space $\mathcal{H}(K)$ by

(4)
$$\langle f,g \rangle_{\mathcal{H}(K)} = \int_{\mathbb{R}^d} \frac{\hat{f}(\boldsymbol{\omega})\overline{\hat{g}(\boldsymbol{\omega})}}{\hat{K}(\boldsymbol{\omega})} \mathrm{d}\boldsymbol{\omega}, \ f,g \in \mathcal{H}(K),$$

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up to a constant factor. Here, \overline{g} denotes the complex conjugate of a function g. Note how the spectrum of K penalizes the spectrum of the functions in $\mathcal{H}(K)$. Roughly, the Hilbert space $\mathcal{H}(K)$ consists of functions f for which $\hat{f}/\sqrt{\hat{K}}$ is square integrable over \mathbb{R}^d . The subtle connections of the Hilbert space $\mathcal{H}(K)$ to sample paths of Gaussian processes with covariance function K will come up at many places in this paper, e.g. in Sections 2, 4.4, 6.3, and 7.1. In this sense, kernels are important links between deterministic and probabilistic models.

A strictly positive definite kernel K is called *stationary* if $K(\boldsymbol{x}, \boldsymbol{y}) \equiv K(\boldsymbol{x} - \boldsymbol{y})$. According to Bochner's theorem [28], K is the Fourier transform of a positive and bounded measure F, that is

$$K(\boldsymbol{x} - \boldsymbol{y}) = \int_{\mathbb{R}^d} \mathrm{e}^{\mathrm{i}(\boldsymbol{x} - \boldsymbol{y},\, \boldsymbol{\omega})} F(\mathrm{d}\boldsymbol{\omega}), \qquad \boldsymbol{x}, \boldsymbol{y} \in \mathbb{R}^d.$$

Here, (\cdot, \cdot) is the inner product in \mathbb{R}^d and i is the unit complex number. Fourier inversion is possible when Kis absolutely integrable, in which case we call the Fourier transform \widehat{K} its spectral density. We note that \widehat{K} is nonnegative and integrable. Furthermore, most of the paper assumes stationarity and isotropy for the covariance function, K, so that

(5)
$$\operatorname{Cov}(Z(\boldsymbol{x}), Z(\boldsymbol{y})) = K(\boldsymbol{x} - \boldsymbol{y}) = \sigma^2 \varphi(\|\boldsymbol{x} - \boldsymbol{y}\|),$$

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for $x, y \in \mathbb{R}^d$ and $\|\cdot\|$ denoting the Euclidean distance. Here, we assume φ to be continuous with $\varphi(0) = 1$. Throughout, we shall equivalently call φ a *function* or a *correlation function*, the last as a shortcut to $\varphi(\|\cdot\|)$. Hence, the parameter $\sigma^2 > 0$ is the variance of Z(x), for all $x \in \mathbb{R}^{d}$. Let Φ_d denote the class of such functions φ inducing a covariance function K through the iden- 100 tity (5) *i.e.* Φ_d is the class of continuous isotropic cor- 101 relation functions defined on \mathbb{R}^d . Such functions have a 102 precise integral representation according to Schoenberg [142], given by

(6)
$$\varphi(x) = \int_0^\infty \Omega_d(rx) F_d(\mathrm{d}r), \qquad x \ge 0,$$

with F_d being a probability measure and

(7)
$$\Omega_d(x) = \Gamma(d/2) \left(\frac{2}{x}\right)^{d/2-1} J_{d/2-1}(x), \qquad x \ge 0,$$

with $\Gamma(\cdot)$ the gamma function and J_{ν} the Bessel function of the first kind of order $\nu > 0$ [118, formula 10.2.2]. For 105 a member φ of the class Φ_d , we can use that its d-variate 106 Fourier transform of $\varphi(||\boldsymbol{x} - \boldsymbol{y}||)$ is isotropic again, and 107 therefore reducible to a scalar integral formula 108 (8)

$$\widehat{\varphi}(z) = \frac{z^{1-d/2}}{(2\pi)^{d/2}} \int_0^\infty u^{d/2} J_{d/2-1}(uz)\varphi(u) \mathrm{d}u, \ z \ge 0,$$

defining its *d*-variate *isotropic spectral density*, and we ¹¹² 84 assume this integral to exist. If the denominator $(2\pi)^{d/2}$ 113 85 is omitted, the same formula holds for the inverse ra- 114 86 dial Fourier transform. Throughout, we write Φ_∞ for $^{\scriptscriptstyle 115}$ 87 $\bigcap_{d>1} \Phi_d$, the class of functions φ inducing positive def-¹¹⁶ 88 inite radial functions on every d-dimensional Euclidean 117 89 118 space. Hence, $\varphi \in \Phi_d$ if and only if $\varphi(\|\cdot\|)$ is a correla-90 tion function in \mathbb{R}^d . 91

2. THE MATÉRN MODEL

The *Matérn model*, $\mathcal{M}_{\nu,\alpha}$, is defined as [148]

(9)
$$\mathcal{M}_{\nu,\alpha}(x) = \frac{2^{1-\nu}}{\Gamma(\nu)} \left(\frac{x}{\alpha}\right)^{\nu} \mathcal{K}_{\nu}\left(\frac{x}{\alpha}\right), \qquad x \ge 0$$

with $\alpha > 0$ the *scale* parameter, $\nu > 0$ the *smoothness* pa-92 rameters, and \mathcal{K}_{ν} a modified Bessel function of the sec-93 ond kind of order ν [2, 9.6.22]. It can be verified that 94 $\mathcal{M}_{\nu,\alpha}(0) = 1$, so that (9) is a correlation function. Argu-95 ments in Stein [148, p48] show that $\mathcal{M}_{\nu,\alpha}$ belongs to the 119 96 class Φ_{∞} . The function $\sigma^2 \mathcal{M}_{\nu,\alpha}$ will be termed *Matérn* 120 97 *covariance function*, and $\sigma^2 > 0$ will denote the variance ₁₂₁ 98 of the associated Gaussian random field. 99

The importance of the Matérn class stems from the parameter ν that controls the differentiability of the sample paths of the associated Gaussian field. Specifically, 125 for any positive integer k, the sample paths of a Gaus- 126 sian field Z on \mathbb{R}^d with Matérn correlation function are 127

k-times mean square differentiable (in any direction) if and only if $\nu > k$. Also, a rescaled version of the Matérn correlation function converges to the Gaussian or squared exponential kernel as $\nu \to \infty$, that is

(10)
$$\mathcal{M}_{\nu,\alpha/(2\sqrt{\nu})}(x) \xrightarrow[\nu \to \infty]{} \exp(-x^2/\alpha^2), \qquad x \ge 0,$$

with convergence being uniform on any compact set of \mathbb{R}^d . For this reason, the parametrisation $\mathcal{M}_{\nu,\alpha/(2\sqrt{\nu})}$ is sometimes also adopted [169].

When $\nu = k + 1/2$, for k a nonnegative integer, the Matérn correlation function simplifies into the product of a negative exponential correlation function with a polynomial of order k. For instance, $\mathcal{M}_{1/2,1}(x) = \exp(-x)$ and $\mathcal{M}_{3/2,1}(x) = \exp(-x)(1+x)$. In general, (11)

$$\mathcal{M}_{k+1/2,1}(x) = \exp(-x) \sum_{i=0}^{k} \frac{(k+i)!}{2k!} \binom{k}{i} (2x)^{k-i}$$

for $k \in \mathbb{N}_0$. This simple algebraic form for the Matérn correlation functions has undoubtedly contributed to the widespread popularity of the Matérn model.

Now we are in a position to explore in detail the many faces of the Matérn model. Section 3 discusses maximum likelihood estimation, Kriging prediction, and the screening effect, while Section 4 explores an SPDE characterisation of the Matérn model. Section 5 discusses the Matérn model as a building block to more sophisticated models, while Section 6 views the scientific landscape through the lens of the Matérn model, with special emphasis on numerical analysis, probability theory and machine learning. Section 7 introduces some recently developed alternatives and generalisations of the Matérn model, while Section 8 compares these alternative models in terms of estimation, prediction, and the screening effect.

3. ESTIMATION AND PREDICTION WITH THE MATÉRN MODEL

Let $D \subset \mathbb{R}^d$ be a subset of \mathbb{R}^d . Consider a set $X_n =$ $\{x_1, \ldots, x_n\}$ of (distinct) locations in D, at which values $\mathbf{Z}_n = (Z(\mathbf{x}_1), \dots, Z(\mathbf{x}_n))^\top$ of the Gaussian random field Z, defined in Section 1.1, are observed. An important problem concerns the *prediction* of values $Z(x_0)$ at an unobserved location $x_0 \in D \setminus X_n$. Then an especially natural predictor for $Z(\boldsymbol{x}_0)$ is

(12)
$$\widehat{Z}_n = \boldsymbol{c}_n^\top \boldsymbol{R}_n^{-1} \boldsymbol{Z}_n$$

with the vector $[\boldsymbol{c}_n]_i = K(\boldsymbol{x}_0, \boldsymbol{x}_i)$ and the kernel matrix $[\mathbf{R}_n]_{i,j} = K(\mathbf{x}_i, \mathbf{x}_j)$. The predictor (12) can be motivated from multiple directions. Classically, (12) is motivated as the best linear unbiased predictor (BLUP) for $Z(x_0)$, and is often referred to as the simple Kriging predictor of $Z(x_0)$ [44]. From a modern perspective, where the role of unbiased estimation is increasingly questioned, we can motivate this choice using alternative optimality properties, including:

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- 1. it is the expectation of $Z(x_0)$ conditionally on the 173 realisation Z_n ;
- 2. it is the optimal estimate (i.e. the Bayes act) for 175 130 $Z(\boldsymbol{x}_0)$ based on the data-set \boldsymbol{Z}_n , under squared error loss [116, Section 13.3]; 132 177

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- 3. it yields the minimal RKHS norm interpolant of the 178 133 data evaluated at x_0 , by Section 6.1; 179
- 4. it is the algorithm for approximating $Z(\boldsymbol{x}_0)$ from 180 135 \boldsymbol{Z}_n that minimises the worst case error in the 136 sense of information-based complexity [116, Sec-137 tion 10.2] and approximation theory (see Section 138 6.1). 139

to name but a few. The Matérn model provides a natu-140 ral setting to study the performance of (12) if we sup-141 pose Z to have a stationary isotropic covariance function 142 $\sigma^2 \mathcal{M}_{\nu,\alpha}$. The crucial question of how to select suitable 143 values for the parameters σ , α , ν will be considered first, 144 in Section 3.1, and then the performance of (12) will be 145 studied in Section 3.2. The possibility of a direct exten-146 sion of the Matérn model to more general domains, such 147 as manifolds and graphs, is discussed in Section 3.3. 148

3.1 Estimation Using Maximum Likelihood 149

Maximum likelihood (ML) and similar estimation 150 methods are popular in this setting due to the availabil-151 ity of practical (inc. gradient-based) numerical methods 152 184 for computation and the classical theory that underpins 153 185 ML. On the other hand, implicit in the use of ML is that 154 186 the statistical model is well-specified, and this judgement 155 must be made on a case-by-case basis. To limit scope, 156 we focus on ML estimation in the sequel. Our aim is to 157 understand when the parameters of the Matérn model can 158 be consistently estimated from data, and to understand the 159 191 asymptotic distribution of the ML estimator. To this end, 160 recall that the Gaussian log-likelihood function is 161

(13)
$$\mathcal{L}_n(\boldsymbol{\theta}) = -\frac{1}{2} \left(\log(|\sigma^2 \boldsymbol{R}_n)|) + \frac{1}{\sigma^2} \boldsymbol{Z}_n^\top \boldsymbol{R}_n^{-1} \boldsymbol{Z}_n \right),$$

up to an additive constant, with $\theta = (\nu, \alpha, \sigma^2)$. The ML estimator is defined as

(14)
$$\widehat{\boldsymbol{\theta}}_n = \operatorname*{argmax}_{\boldsymbol{\theta} \in \mathbb{R}^3_+} \mathcal{L}_n(\boldsymbol{\theta}).$$

The ML estimate for the variance parameter can be com-162 puted in closed-form as $\hat{\sigma}_n^2 = \mathbf{Z}_n^{\uparrow} \mathbf{R}_n^{-1} \mathbf{Z}_n / n$; plugging 163 this expression into (13) reduces the numerical problem to 164 optimisation of a so-called concentrated likelihood over 165 \mathbb{R}^2_+ . However, maximizing the log-(concentrated) like-166 lihood requires a nonlinear optimisation problem to be 167 solved, for which numerical methods must be used; see 168 Section 4.3. 169

The performance of ML estimation has been studied 170 principally in two different asymptotic limits. Under fixed 171 domain asymptotics, the sampling domain D is bounded 172

and the set of sampled locations X_n becomes increasingly dense in D. Under increasing domain asymptotics, the domain D grows with the number n of observed data, and the distance between any two sampled locations is bounded away from zero. Zhang and Zimmerman [180] note that the peformance of the ML estimator can be quite different under these two frameworks, as will now be discussed.

3.1.1 Increasing Domain Asymptotics. Mardia and Marshall [107] make use of increasing domain asymptotics to establish, under mild regularity conditions, that the ML estimator is strongly consistent, meaning that $\widehat{\boldsymbol{\theta}}_n \stackrel{a.s.}{\longrightarrow} \boldsymbol{\theta}_0$ for the *true* parameter $\boldsymbol{\psi}_0$. Furthermore, they establish that the ML estimator is asymptotically normal, meaning that

(15)
$$\boldsymbol{F}^{1/2}(\boldsymbol{\theta}_0)(\widehat{\boldsymbol{\theta}}_n - \boldsymbol{\theta}_0) \xrightarrow{d} \mathcal{N}(\boldsymbol{0}, \boldsymbol{I})$$

where $F(\theta) = -E[\mathcal{L}''_n(\theta)]$ is the Fisher information matrix, whose entries are

$$F(\boldsymbol{\theta})_{i,j} = \frac{1}{2} \operatorname{tr} \left(\frac{\mathrm{d} \boldsymbol{\Sigma}_n}{\mathrm{d} \boldsymbol{\theta}_i} \boldsymbol{\Sigma}_n^{-1} \frac{\mathrm{d} \boldsymbol{\Sigma}_n}{\mathrm{d} \boldsymbol{\theta}_j} \boldsymbol{\Sigma}_n^{-1} \right),$$

and $\boldsymbol{\Sigma}_n = \sigma^2 \boldsymbol{R}_n$. Although our focus is on the Matérn model, we note that these kind of asymptotic results hold for any parametric correlation function obeying particular regularity conditions that are stated in terms of eigenvalue conditions on the correlation matrix and its derivatives [107], thought these may not be easy to verify in general (see for instance Shaby and Ruppert [144], for the exponential case). Generally speaking, as long as the spatial extent of the sampling region is large compared with the range of dependence of the random field, increasingdomain asymptotics provide a very accurate description of the behavior of the ML estimate [180, 144, 84].

3.1.2 Fixed Domain Asymptotics. Zhang [179] considered ML estimation for the Matérn model under fixed domain asymptotics, proving that when the smoothness parameter ν is known and fixed, none of the parameters σ^2 and α can be estimated consistently when d = 1, 2, 3. Instead, only the parameter

(16)
$$\operatorname{micro}_{\mathcal{M}} = \sigma^2 / \alpha^{2\nu},$$

sometimes called *microergodic* parameter [180, 148], can be consistently estimated. This is a consequence of the equivalence of the two corresponding Gaussian measures, that we denote with $P(\sigma_i^2 \mathcal{M}_{\nu,\alpha_i})$, with i = 0, 1. In particular, for any bounded infinite set $D \subset \mathbb{R}^d$, d = 1, 2, 3, $P(\sigma_0^2 \mathcal{M}_{\nu,\alpha_0})$ is equivalent to $P(\sigma_1^2 \mathcal{M}_{\nu,\alpha_1})$ on the paths of $Z(\boldsymbol{x}), \boldsymbol{x} \in D$, if and only if

(17)
$$\sigma_0^2/\alpha_0^{2\nu} = \sigma_1^2/\alpha_1^{2\nu}.$$

In contrast, for $d \ge 5$, Anderes [8] proved the orthogonal-193 ity of two Gaussian measures with different Matérn co-194 variance functions and hence, in this case, all the param-195 eters can be consistently estimated under fixed-domain 196 asymptotics. The case d = 4 has been recently studied in 197 Bolin and Kirchner [31]. 198

Asymptotic results associated with ML estimation of the microergodic parameter, again for a fixed known smoothness parameter ν , can be found in Zhang [179], and later on in Kaufman and Shaby [84]. In particular, for a zero mean Gaussian field defined on a bounded infinite set $D \subset \mathbb{R}^d$, d = 1, 2, 3, with a Matérn covariance function $\sigma_0^2 \mathcal{M}_{\nu,\alpha_0}$ the ML estimator $\hat{\sigma}_n^2 / \hat{\alpha}_n^{2\nu}$ of the microergodic parameter is strongly consistent, *i.e.*,

$$\hat{\sigma}_n^2 / \hat{\alpha}_n^{2\nu} \xrightarrow{a.s.} \sigma_0^2 / \alpha_0^{2\nu},$$

and its asymptotic distribution is given by

$$\sqrt{n}(\hat{\sigma}_n^2/\hat{\alpha}_n^{2\nu} - \sigma_0^2/\alpha_0^{2\nu}) \stackrel{d}{\longrightarrow} \mathcal{N}(0, 2(\sigma_0^2/\alpha_0^{2\nu})^2)$$

Generally speaking, when the range of dependence of the 199 random field is large with respect to the spatial extent of 200 the sampling region, fixed domain asymptotics provide a 201 very accurate description of the behavior of the ML es-202 timate of the microergodic parameter [84]. Extensions of 203 these results to the case where Z is observed with Gaus-204 sian errors can be found in Tang et al. [156], while re-205 sults for a space-time version of the Matérn model can be 206 found in Ip and Li [77] and Faouzi et al. [55]. Finally we ²²⁵ 207 highlight that the efficient estimation of the microergodic 226 208 parameter assuming the smoothness parameter unknown ²²⁷ 209 is still an open problem; some promising results in this ²²⁸ 210 direction can be found in Loh et al. [105]. 211

3.2 Prediction and the Screening Effect 212

The equivalence of Gaussian measures within the 232 213 Matérn class has consequences for prediction of $Z(x_0)$ at ²³³ 214 an unobserved location $oldsymbol{x}_0 \in D \setminus X_n$; these consequences 234 215 will now be discussed. In what follows, ν is supposed ²³⁵ 216 known and fixed, and we consider the setting where σ ²³⁶ 217 and α are *misspecified*. That is, we suppose Z is a Gaus-237 218 sian field with Matérn covariance $\sigma_0^2 \mathcal{M}_{\nu,\alpha_0}$, and we con-238 219 sider the performance of the predictor (12) when a Matérn ²³⁹ 220 model $\sigma_1^2 \mathcal{M}_{\nu,\alpha_1}$ is used. This situation is typical, since the 240 221 true parameters σ_0 and α_0 of the data-generating process 222 241 will be unknown in general. Our theoretical setting will 223 242 be fixed domain asymptotics. 224 243

Note, first, that (12) does not depend on the value of σ_1 , but does depend on the value of the parameter α_1 (and the parameter ν , but this parameter is fixed). This depen-246 dence will be emphasised using the notation $c_n(\alpha_1)$ and 247 $\boldsymbol{R}_n(\alpha_1)$. Under the Gaussian measure $P(\sigma_0^2 \mathcal{M}_{\nu,\alpha_0})$ associated with the *true* model $\sigma_0^2 \mathcal{M}_{\nu,\alpha_0}$, the mean squared

error of the predictor $\widehat{Z}_n(\alpha_1)$ is given by

$$\begin{aligned} \operatorname{VAR}_{\alpha_0,\sigma_0^2} \left[\widehat{Z}_n(\alpha_1) - Z(\boldsymbol{x}_0) \right] \\ &= \sigma_0^2 \Big(1 - 2\boldsymbol{c}_n(\alpha_1)^\top \boldsymbol{R}_n(\alpha_1)^{-1} \boldsymbol{c}_n(\alpha_0) \\ &+ \boldsymbol{c}_n(\alpha_1)^\top \boldsymbol{R}_n(\alpha_1)^{-1} \boldsymbol{R}_n(\alpha_0) \boldsymbol{R}_n(\alpha_1)^{-1} \boldsymbol{c}_n(\alpha_1) \Big), \end{aligned}$$

while if there is no misspecification then the previous expression reduces to

(18)
$$\operatorname{VAR}_{\alpha_0,\sigma_0^2} \left[\widehat{Z}_n(\alpha_0) - Z(\boldsymbol{x}_0) \right]$$
$$= \sigma_0^2 \left(1 - \boldsymbol{c}_n(\alpha_0)^\top \boldsymbol{R}_n^{-1}(\alpha_0) \boldsymbol{c}_n(\alpha_0) \right).$$

Under regularity conditions, and for fixed domain asymptotics, Stein [146] shows that both asymptotically efficient prediction and asymptotically correct estimation of prediction variance hold when the two Gaussian measures $P(\sigma_i^2 \mathcal{M}_{\nu,\alpha_i}), i = 0, 1$ are equivalent, *i.e.* (17). Specifically,

(19)
$$\frac{\operatorname{VAR}_{\sigma_0^2,\alpha_0} [\widehat{Z}_n(\alpha_1) - Z(\boldsymbol{x}_0)]}{\operatorname{VAR}_{\sigma_0^2,\alpha_0} [\widehat{Z}_n(\alpha_0) - Z(\boldsymbol{x}_0)]} \xrightarrow{a.s.} 1$$

and

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(20)
$$\frac{\operatorname{VAR}_{\sigma_{1}^{2},\alpha_{1}}\left[\widehat{Z}_{n}(\alpha_{1})-Z(\boldsymbol{x}_{0})\right]}{\operatorname{VAR}_{\sigma_{0}^{2},\alpha_{0}}\left[\widehat{Z}_{n}(\alpha_{1})-Z(\boldsymbol{x}_{0})\right]} \xrightarrow{a.s.} 1$$

The implication of (19) is that, under the true model, if the correct value of ν is used, any value of α_1 will give asymptotic efficiency. The implication of (20) is stronger and guarantees that using the misspecified predictor under the correct and misspecified models is asymptotically equivalent from mean squared error point of view. Note that these kind of results does not consider the uncertainty associated with the covariance parameters of the misspecified model. Kaufman and Shaby [84] show that (20) still holds by considering the ML estimator of the variance $\hat{\sigma}_n^2 = \mathbf{Z}_n^{\top} \mathbf{R}_n^{-1}(\alpha_1) \mathbf{Z}_n / n$ in place σ_1^2 .

Conditions of equivalence of two Gaussian measures based on a space-time [77] and bivariate [14] version of the Matérn model have also been established. Next, we consider a practically important aspect of prediction; the co-called screening effect.

Screening Effect. The screening effect refers to the phenomenon where the predictor (12) depends almost exclusively on those observations that are located nearest to the predictand [149]. As such, the screening effect is an important tool that can be used to mitigate the computational burden of evaluating (12) in the presence of big datasets. This issue has traditionally been an important subject in geostatistics [109, 110, 111, 41]. Indeed, Matheron [109, 110], in the School of Geostatistics at the Ecole des Mines, developed a first formalisation of screening effect, referring to situations where the observa- $_{261}$ tions located far from the predictand receive a zero krig- $_{262}$ ing weight. Matheron's definition has a direct connection $_{263}$ with the Markov property on the real line, which happens $_{264}$ when kriging is performed under the exponential model $_{265}$ (indeed, $\mathcal{M}_{1/2,\alpha}$).

M. Stein [148, 149, 151, 152] adopts an alternative definition of the screening effect that will now be de-²⁶⁷ scribed. Let Z be a mean-square continuous, zero mean ²⁶⁸ and weakly stationary Gaussian random field on \mathbb{R}^d . Let ²⁶⁹ $e(X_n)$ be the error of the predictor (12) of $Z(x_0)$ based on ²⁷⁰ Z_n . Two choices for the set X_n of observation locations ²⁷¹ will be considered, and to this end we let $F_{\epsilon}, N_{\epsilon}$ be sets, ²⁷² indexed by $\epsilon > 0$, such that N_{ϵ} contains the nearest obser-²⁷³ vations to the predictand, and F_{ϵ} the furthest observations. ²⁷⁴ Then Stein [149] says that N_{ϵ} asymptotically screens out F_{ϵ} when ²⁷⁵

(21)
$$\lim_{\epsilon \downarrow 0} \frac{\mathbb{E} e(N_{\epsilon} \cup F_{\epsilon})^2}{\mathbb{E} e(N_{\epsilon})^2} = 1.$$

A thorough discussion of the implications of this definition can be found in Porcu et al. [128], where nontrivial differences between fixed domain and increasing domain asymptotics are reported.

The spatial configuration of the sampling point X_n de-285 termines whether the screening effect will hold. Porcu 286 et al. [128] refer to a *regular scheme* as one for which $F_{\epsilon} = \{\epsilon(\boldsymbol{x}_0 + j)\}, \text{ for } j \in \mathbb{Z}^d \text{ and } N_{\epsilon} \text{ being the restric-}$ 288 tion of F_{ϵ} to some fixed region with x_0 in its interior, 289 assuming $x_0 \notin \mathbb{Z}^d$. For regular schemes, Stein [149] es-290 tablished (21) whenever the spectrum \hat{K} varies regularly 291 at infinity [27] in every direction with a common index of 292 *variation* [quoted from 128]. However, this condition may 293 not be useful for space-time processes, where differentia-294 bility properties in the space and time coordinates are not 295 necessarily identical. To overcome such a problem, we instead consider an *irregular scheme*: for x_1, \ldots, x_n being ₂₉₇ distinct nonzero elements of $\mathbb{R}^d, oldsymbol{y}_1, \dots, oldsymbol{y}_N$ distinct ele- 298 ments of \mathbb{R}^d , $\boldsymbol{x}_0 = \boldsymbol{0} \in \mathbb{R}^d$ and $\boldsymbol{y}_0 \in \mathbb{R}^d$ being nonzero, we 299 have $N_{\epsilon} = \{\epsilon \boldsymbol{x}_1, \dots, \epsilon \boldsymbol{x}_n\}$ and $F_{\epsilon} = \{\boldsymbol{y}_0 + \epsilon \boldsymbol{y}_1, \dots, \boldsymbol{y}_0 +$ $\{\epsilon y_N\}$. The Stein hypothesis [termed in 128]

(22)
$$\forall R > 0$$
, $\lim_{\|\boldsymbol{\omega}\| \to \infty} \sup_{\|\boldsymbol{\tau}\| < R} \left| \frac{\widehat{K}(\boldsymbol{\omega} + \boldsymbol{\tau})}{\widehat{K}(\boldsymbol{\omega})} - 1 \right| = 0$, ³⁰⁰

provides a sufficient condition for the screening effect in ³⁰² this setting (under some mild additional conditions on \hat{K} ³⁰³ and N_{ϵ}), which can be verified in dimensions d = 1 and ³⁰⁴ d = 2 for mean-square continuous but non-differentiable ³⁰⁵ random fields, for some specific designs N_{ϵ} [151]. The Matérn model with $K = \mathcal{M}_{\alpha,\nu}$ admits a simple expression for its spectrum [2, 11.4.44]:

(23)
$$\widehat{\mathcal{M}}_{\nu,\alpha}(z) = \frac{\Gamma(\nu + d/2)}{\pi^{d/2}\Gamma(\nu)} \frac{\alpha^d}{(1 + \alpha^2 z^2)^{\nu + d/2}}, \ z \ge 0,$$

from which (22) can be verified.

The screening effect can thus be established for the Matérn model, under both regular and irregular schemes, justifying the use of "local" approximations to the predictor (12).

3.3 Matérn on Manifolds and Graphs

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Let M be a general manifold. A pragmatic question is whether the Matérn correlation function (9) can be composed with a suitable metric q, defined on the manifold, to preserve positive definiteness over M. For the case of the sphere, a natural metric is the geodesic distance; the length of the arc connecting any pair of points located over the spherical shell. For this metric, $(x, y) \mapsto$ $\mathcal{M}_{\nu,\alpha}(q(x,y))$ is a correlation function only for $0 < \nu \leq 1$ 1/2 [63]. This limitation is emphasised in Alegría et al. [3], who propose the \mathcal{F} family, a model that is valid on the sphere, and having the same properties as the Matérn function in terms of mean-square differentiability and fractal dimension. The Matérn function on other general manifolds has been studied by Li et al. [98]. Guinness and Fuentes [69] propose a spectral expansion to define a covariance function that mimics the Matérn model, but this construction is criticised by Lindgren et al. [101] as being incorrect as the spectral expansion does not reproduce the same properties of the Matérn model.

Unfortunately, it seems that the limited applicability of the Matérn model on any space that is not a flat surface extends to more abstract settings as well. An elegant isometric embedding argument in Anderes et al. [9] proves that the restriction $0 < \nu \le 1/2$ is required when the input space is a graph with Euclidean edges. A more general argument in Menegatto et al. [112] proves that the same restriction is inherited for a general quasi metric space endowed with a geodesic metric. The notable effort by Bolin and Kirchner [30] provides a model that is once differentiable over metric graphs. It is reasonable to conclude that some form of the SPDE approach, which we discuss next in Section 4.2, is needed in general to extend the Matérn model to a general manifold.

4. THE MATÉRN MODEL IN COMPUTATIONAL STATISTICS

This section explores the interaction of the Matérn model with computational statistics, starting with numerical methods for *implementation* of the Matérn model (Sections 4.1, 4.2 and 4.3), and then turning to uses of the Matérn model to *facilitate* numerical computation itself (Section 4.4).

4.1 Implementation as an SDE

The Matérn model admits a *state space* representation as an SDE, which enables efficient computational techniques from the signal processing literature to be employed for simulation, estimation and prediction. Indeed,

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focusing on dimension d = 1, and letting

$$\mathbf{Z}(x) = (Z, \mathrm{d}Z/\mathrm{d}x, \dots, \mathrm{d}^k Z/\mathrm{d}x^k),$$

the Matérn model $\mathcal{M}_{\nu,\alpha}$ with $\nu=k+1/2$ admits the 336 characterisation

$$d\mathbf{Z} = \begin{pmatrix} 0 & 1 \\ \vdots & \vdots \\ 0 & 1 \\ -a_0 - a_1 \dots - a_{k-1} \end{pmatrix} \mathbf{Z} \, dx + \begin{pmatrix} 0 \\ \vdots \\ 0 \\ 1 \end{pmatrix} d\mathcal{W}$$

343 where $a_i = {}_{k+1}C_i \cdot \alpha^{-k-1+i}$, the .C. are binomial coef-307 ficients, and $\mathcal{W}(x)$ represents a zero-mean white noise 308 process on $x \in \mathbb{R}$ [73]. The advantage of state space for-309 346 mulations is that both estimation and prediction can be 310 347 performed in a single pass through the data, at linear 311 348 O(n) cost, using familiar Kalman updating equations as 312 349 described in Sarkka et al. [135] and in further detail in 313 350 Chapter 6 of Hennig et al. [75]. Similar characterisations 314 351 for higher dimensions, including spatio-temporal versions 315 352 of the Matérn model, can be found in Sarkka et al. [135], 316 353 though we note these retain linear complexity only in the 317 354 number of time steps; complexity is cubic in the size of 318 355 the spatial grid. The SPDE approach can offer a solution 319 356 in this respect, and we discuss this next. 320 357

321 4.2 Implementation as an SPDE

A major reason for the continued popularity of the Matérn model is the availability of efficient and scalable numerical methods for simulation, due in large part to Lindgren et al. [102]. These authors consider the SPDE

(24)
$$(\alpha^{-2} - \Delta)^{\gamma/2} Z(\boldsymbol{x}) = \mathcal{W}(\boldsymbol{x}), \qquad \boldsymbol{x} \in \mathbb{R}^d,$$

where $\alpha > 0$, Δ is the Laplacian, and \mathcal{W} is a Gaussian white noise on \mathbb{R}^d , so that $\text{Cov}(\mathcal{W}(A_1), \mathcal{W}(A_2)) = |A_1 \cap A_2|$, where A_i are subsets of \mathbb{R}^d , i = 1, 2, and where $|\cdot|$ is the volume integral. Whittle [167] and Whittle [168] proved that the solution to (24) is a Gaussian field with Matérn covariance $\sigma^2 \mathcal{M}_{\nu,\alpha}$ with parameters α (as before) and

$$\sigma^2 = \frac{\Gamma(\nu)\alpha^{2\nu}}{\Gamma(\nu + d/2)(4\pi)^{d/2}}, \qquad \nu = \gamma - d/2.$$

This perspective offers two insights; first, tools developed 367 322 for the numerical approximation of SPDEs can be brought 368 323 to bear on the Matérn model, and second, there is a clear 369 324 path to generalise the definition of the Matérn model to 370 325 any (planar or non planar) manifold on which the analo- 371 326 gous SPDE may be defined. (For example, Jansson et al. 372 327 [78] take this perspective to generalise the Matérn model 373 328 to the sphere \mathbb{S}^d .) 374 329

To provide a computationally convenient approximation to (24), Lindgren et al. [102] considered the weak solution to (24) and approximation of the weak solution rot (24) and approximation of the weak solution

using basis functions with compact support over a compact domain $\Omega \subset \mathbb{R}^d$ (specifically, a *Galerkin* approximation using finite element basis functions was used). As a result, the authors establish a formal route to approximation of the random field Z with a *Gauss–Markov* random field having a *sparse* precision matrix. Sparse matrix algebra enables fast simulation of realisations from the Matérn random field, and fast evaluation of the likelihood (13) (albeit not fast evaluation of the gradient of the likelihood).

The choice of domain Ω introduces boundary effects which must be carefully mitigated. Khristenko et al. [87], Brown et al. [37] provide a solution for the case where γ is an integer; the non-integer case is considered in Bolin and Kirchner [30]. The extension of the Matérn field based on SPDEs to space-time is provided by Cameletti et al. [39] and subsequently by Bakka et al. [15], Clarotto et al. [42], while the multivariate Matérn case has been explored in Bolin and Wallin [33]. Alternative approximations based on Galerkin methods on manifolds have been provided by Lang and Pereira [92]. An interesting approach that allows working on manifolds with huge datasets is proposed by Pereira et al. [121]. The interest in this literature is dual. On the one hand, the technical aspects related to the finite dimensional representation of Gaussian random fields are extremely interesting per se. On the other hand, this group of authors is actually driven by providing tools for efficient computation. This is witnessed by the relevant existing packages (R-INLA, inlabru, and rSPDE for instance) and we refer to the review of Lindgren et al. [101].

Sanz-Alonso and Yang [133] attempt to explain the trade-off between accuracy and scalability in numerical approxmation of the Matérn model. Recall that, in the SPDE approach [102], Z in (24) is numerically approximated using a Gaussian process

(25)
$$Z_{\delta}(\boldsymbol{x}) = \sum_{k=1}^{n_{\delta}} \omega_k \epsilon_k(\boldsymbol{x}), \qquad \boldsymbol{x} \in \Omega,$$

where ϵ_k are finite element basis functions and the vector $\boldsymbol{\omega} = (\omega_1, \dots, \omega_{n_\delta})^{\top}$ is multivariate Gaussian with zero mean and with a sparse precision matrix. The accuracy of the approximation Z_{δ} is dependent on (a) the compact support of the finite elements basis functions, (b) boundary effects due to the domain Ω , and (c) by the mesh width δ that determines the cardinality n_{δ} in (25). Most of the earlier literature has considered (25) with n_{δ} proportional to the sample size n of the dataset being modelled. Sanz-Alonso and Yang [133] adopt a fixed domain asymptotic approach to explain when $n_{\delta} \ll n$ might be a legitimate strategy. To do so, they consider Gaussian process regression and work under the framework of Bayesian contraction rates. Their results provide justification for specific

scalings of n_{δ} with $n_{\delta} = o(n)$, provided that the smoothness ν is sufficiently high.

A different path to SPDE and Gauss-Markov random 434 380 fields was recently taken in Sanz-Alonso and Yang [134], 435 381 who adopt graph-based discretisations of SPDEs. This ap- 436 382 proach can be well-suited to working with discrete and 437 383 unstructured point clouds, such as in machine learning 438 384 tasks where the data belong to an implicitly defined low-385 439 dimensional manifold. A second advantage of this ap-386 proach is that an explicit triangulation of the domain is 387 not required. 388

4.3 Approximate Likelihood and the Matérn Model

In estimating the parameters of the Matérn model using 390 ML (14), numerical optimisation is required. Although 391 generic optimisation routines can be used, an often better 392 approach is to first construct an accurate-but-cheap ap-393 proximation to the likelihood, which can then be more 394 442 readily maximised. Indeed, approximate likelihoods are 395 essential when dealing with large datasets, since the eval-396 uation of (13) requires computing the inverse and the 397 determinant of the correlation matrix, usually via the 398 446 Cholesky decomposition at complexity $O(n^3)$ and stor-399 447 age cost $O(n^2)$. 400

448 Perhaps the most successful approximation is Vecchia's 401 method [163], which has attracted a remarkable amount 402 of attention in recent times [inc. 147, 48, 49, 67, 47]. The 403 Vecchia approximation can be used with any correlation 451 404 model and its basic idea is is to replace (13) with a prod-452 405 uct of Gaussian conditional distributions, in which each 453 406 conditional distribution involves only a small subset of 454 407 the data. This approximation requires that the data are or-408 *dered* and the number m of 'previous' data on which to 456 409 condition is to be specified. Generally, larger m entails 410 more accurate and computationally expensive approxima-411 tion, while the choice of ordering affects the accuracy of 459 412 the approximation [67]. The Vecchia method provides a ⁴⁶⁰ 413 sparse approximation to the Cholesky factor of the pre- 461 414 cision matrix, such that the approximate likelihood can 462 415 be computed in $O(nm^3)$ time and with $O(nm^2)$ storage 463 416 cost. See the recent review of Katzfuss and Guinness [83] 464 417 for further detail. The Vecchia likelihood can be viewed 465 418 as a specific instance of a more general class of estimation 466 419 methods called quasi- or composite likelihood [103, 162] 467 420 that have been widely used for the estimation of Gaussian 468 421 fields with the Matérn model [52, 26, 13]. 422

An alternative method of mitigating the computational 470 423 burden of ML estimation is *covariance tapering* [59]. The 471 424 basic idea is to multiply the Matérn model with a com- 472 425 pactly supported correlation function, resulting in a 'mod- 473 426 ified' Matérn model with compact support. This induces 474 427 sparseness in the associated covariance matrix, so that al- 475 428 gorithms for sparse matrices can be exploited for a com- 476 429 putationally efficient evaluation of the Cholesky decom- 477 430 position [59]. However, some authors [25, 23] suggest 478 431

that tapering might be an obsolete approach in view of the fact that flexible compactly supported models that include the Matérn model as a special case have been recently proposed; see Section 8. A comprehensive review of the likelihood approximations is beyond the scopes of this paper, so we refer the reader to Sun et al. [154] and Heaton et al. [74] for further detail.

4.4 The Matérn Model for Bayesian Computation

In the last decade there has been increasing interest in the use of kernel methods for solving PDEs. Consider a system

$$\mathcal{A}u = f \qquad \text{in } \Omega$$
$$\mathcal{B}u = g \qquad \text{on } \partial \Omega$$

specified by a differential equation involving \mathcal{A} and f, and initial or boundary conditions specified by \mathcal{B} and q. Dating back at least to Fasshauer [57] in the deterministic setting, and reinterpreted through a Bayesian lens by authors such as Cockayne et al. [43], one can seek an approximation to the strong solution $u: \Omega \to \mathbb{R}$ by modelling u as a priori a Gaussian random field and conditioning that field to satisfy the differential equation at locations $\{x_1, \ldots, x_m\} \subset \Omega$ and satisfy the boundary conditions at locations $\{x_{m+1},\ldots,x_n\} \subset \partial \Omega$. The conditional mean of this process coincides with the symmetric collocation method introduced by Fasshauer [57], which we return to in Section 6.1, while the conditional variance provides probabilistic uncertainty quantification for the solution, expressing the uncertainty that remains as a result of using only a finite computational budget. To implement these methods, one requires a Gaussian process whose sample paths possess sufficient regularity for the operation of conditioning on the derivative $\mathcal{A}u$ to be welldefined. On the other hand, assuming excessive smoothness could lead to over-confident uncertainty quantification. One therefore requires a kernel with customisable smoothness, which can be adapted to the differential equation at hand. The Matérn class satisfies this requirement, but is not alone in doing so; we continue discussion of this point in Section 7.

A specific PDE that has received considerable recent attention in the Bayesian statistical community is the *Stein equation*, for which $Au = c + p^{-1}\nabla \cdot (p\nabla u)$, where pis the probability density function of a posterior distribution of interest, f is a function whose posterior expectation we seek to compute, and c is a constant. If the Stein equation has a solution, then c must be the value of the posterior expectation we seek. This has motivated several efforts to numerically solve the Stein equation, as a more direct alternative to first approximating p (for example using Markov chain Monte Carlo) and then using the approximation of p to approximate the expectation of interest. In this context kernel methods are typically

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used [117, 145] and in particular the kernel should have 530 479 smoothness that is two orders higher than that of the func- 531 480 tion f whose expectation is of interest, since the Stein 532 481 equation is a second-order PDE. The generalisation of the 533 482 Stein equation to Riemannian manifolds was considered 534 483 in [19], who advocated for the use of kernels with cus- 535 484 tomisable smoothness that reproduce Sobolev spaces of 536 485 functions on the manifold, such as the (manifold gener- 537 486 alisation of the) Matérn model. The connection between 538 487 the Matérn model and Sobolev spaces is set out in Section 539 488 6.1. 489 540

5. FLEXIBLE MODELLING WITH MATÉRN

543 One might object that the Matérn model is insufficiently 490 flexible for many statistical applications, being limited to 491 scalar-valued random fields that are stationary, isotropic 492 and Gaussian. However, the Matérn model is also an im-493 547 portant building block for many more sophisticated mod-494 els, some of which will now be described. This is a rich 495 549 literature, and our discussion is necessarily succinct; an 496 extended version of this section can be found in Appendix 497 551 A of the Supplementary Material. 498 552

499 5.1 Scalar Valued Random Fields

Let us start by discussing models for scalar-valued random fields that build on the Matérn model. Note that one can trivially introduce non-zero mean functions into the Matérn model, or combine (additively or multiplicatively) kernels to obtain a potentially more expressive kernel; we will not dwell on either point.

560 To relax the isotropy assumption of the Matérn model, 506 561 [7] consider scale mixtures that take into account pref-507 erential directions in which spatial dependence develops. 508 563 On the other hand, the case of space-time models re-509 564 quires special treatment, and non-separable versions of 510 565 the Matérn kernel are described in Gneiting [62], Zas-511 tavnyi and Porcu [178]. 512

567 The stationarity assumption was relaxed in a paramet-513 ric manner in Paciorek and Schervish [119], and then 514 in a nonparametric manner in Roininen et al. [132]. An 569 515 attempt to strike a balance between the computational 516 tractability of parametric models and the flexibility of 517 nonparametric models was reported in Wilson et al. [170], 518 who proposed input warping to transform the inputs to the 519 Matérn model using a neural network. 520

The Gaussian assumption can be relaxed through out-521 *put warping*, meaning transformation of the form Z(x) =522 $w(Z(\boldsymbol{x}))$ where $w(\cdot)$ is a nonlinear map from \mathbb{R}^d to \mathbb{R}^d . 523 The covariance function of \tilde{Z} will not be Matérn in gen-524 eral, when the covariance function of Z is Matérn, but 525 if w is sufficiently regular then the smoothness proper-526 ties of Z transfer to Z. The question of whether there ex-527 ist non-Gaussian processes whose covariance function is 571 528 nevertheless of Matérn class was answered positively in 529

Åberg and Podgórski [1]. Yan and Genton [174] have proposed *trans-Gaussian* random fields with Matérn covariance function. Bolin [29] and subsequently Wallin and Bolin [164] provided SPDE-based constructions for non-Gaussian Matérn fields. General classes of non-Gaussian fields with covariance $g(\mathcal{M}_{\nu,\alpha})$, for $g(\cdot)$ a suitable function that preserves the positive definiteness and smoothness properties of the Matérn model, have been provided for instance by Palacios and Steel [120], Xua and Genton [173], Bevilacqua et al. [24], Morales-Navarrete et al. [113].

An important extension of the Matérn model, which has received recent attention, is to random fields on spaces for which classical notions of smoothness are not well-defined. For example, Anderes et al. [9] consider graphs with Euclidean edges, equipped with either the geodesic distance over the graph, or the resistance metric. Menegatto et al. [112] provide a generalisation of this setting by considering quasi-metric spaces. Bolin et al. [32] adopt a different approach to build random fields with their covariance structure on metric graphs. Space-time version of the Matérn model, for graphs with Euclidean edges, have been considered by Tang and Zimmerman [155] and Porcu et al. [127]. These efforts considerably extend the applicability of the Matérn model.

The Matérn covariance function decays exponentially with distance, which can be inappropriate for modelling processes that involve long memory. Several approaches have been developed to modify the tails of the Matérn correlation function while preserving many of its desirable characteristics; we describe these in Section 7.

[68] considers Gaussian random fields defined for lattices \mathbb{Z}^d with a covariance function that is the restriction of the Matérn covariance to \mathbb{Z}^d . The resulting spectrum is smoothed version of the spectral density associated with the Matérn covariance. For this specific situation, the SPDE approximation can overestimate the scale, α . Yet, it is not clear how this message extends to Gaussian fields that are continuously indexed in \mathbb{R}^d .

5.2 Vector-Valued Random Fields

There has been a plethora of approaches related to multivariate spatial modeling, and the reader is referred to Genton and Kleiber [60]. Here, the isotropic covariance function $\boldsymbol{K} : [0, \infty) \to \mathbb{R}^{p \times p}$ is matrix-valued. The elements on the diagonal, K_{ii} , are called *auto-covariance* functions, and the elements K_{ij} , $i \neq j$, are called *crosscovariance* functions. Gneiting et al. [64] proposed a multivariate Matérn model

(26)
$$K_{ij}(x) = \sigma_{ii}\sigma_{jj}\rho_{ij}\mathcal{M}_{\nu_{ij},\alpha_{ij}}(x), \qquad x \ge 0,$$

where σ_{ii}^2 is the variance of Z_i , the *i*th component of a multivariate random field in \mathbb{R}^p , and ρ_{ij} is the collocated

correlation coefficient. There are restrictions on the pa-572 rameters ν_{ij} , α_{ij} and ρ_{ij} required to ensure positive defi-573 niteness, and often the restrictions on the collocated cor-574 relations coefficients ρ_{ij} are rather strict. This last remark 575 has motivated alternative approaches, and the reader is re-576 ferred to Apanasovich et al. [11] and more recently to 577 Emery et al. [54]. Extensions to multivariate space-time 578 Matérn structures have been provided by Allard et al. 579 [6] and through a technical approach by Porcu et al. 580 [125]. Multivariate nonstationary Matérn functions have 581 been proposed by Kleiber and Nychka [88]. Multivariate 582 Matérn models with *dimple* effect have been studied by 583 Alegría et al. [4]; a 'dimple' in a space-time covariance 584 model refers to the case when Cov(Z(x,t), Z(x',t')) is 585 bigger than $Cov((Z(\boldsymbol{x},t),Z(\boldsymbol{x}',t)))$, which requires spe-586 cial mathematical treatment. 587

Multivariate Matérn modeling on graphs has been re-588 cently investigated in Dey et al. [51], who propose a class 589 of multivariate graphical Gaussian processes through 590 stitching, a construction that gets multivariate covari-591 ance functions from the graph, and ensures process-level 592 conditional independence between variables. When cou-593 pled with the Matérn model, stitching yields a multi-594 variate Gaussian process whose univariate components 595 are Matérn Gaussian processes, and which agrees with 596 process-level conditional independence as specified by 597 the graphical model. Stitching can offer massive com-598 putational gains and parameter dimension reduction. An 599 ingenious approach to Gaussian process construction in-600 volving the Matérn covariance function has been recently 601 proposed by Li et al. [97], who considered a product space 602 involving the d-dimensional Euclidean space cross an ab-603 stract set that allows to index group labels. 604

5.3 Directions, Shapes and Curves

The Matérn model has an important role in the study 606 of directional processes, with Banerjee et al. [18] formal-607 634 ising the notions of directional *finite difference processes* 608 and directional derivative processes with special empha-609 sis on the Matérn model. The Matérn model also has a role 610 in shape analysis, where Banerjee and Gelfand [16] in-611 troduced *Bayesian wombling* to measure *spatial* gradients 612 related to curves through 'wombling' boundaries, and ap-613 proach taken further in Halder et al. [71]. The smoothness 614 properties of the Matérn model are ideally suited to such 615 a framework. Modeling approaches to temporal gradients 616 using the Matérn model have been proposed by Quick 617 et al. [131]. Related to these approaches, the smoothness 618 parameter ν of the Matérn model plays a central role in 619 the recent paper by Halder et al. [71], who analyse ran-620 dom surfaces in order to explain latent dependence within 621 a response variable of interest. 622

This represents a short tour of *statistical* applications of the Matérn model, but its reach goes well beyond statistics, and we explore the importance of the Matérn model to related fields next.

6. THE MATÉRN MODEL OUTSIDE STATISTICS

This section explores the impact of the Matérn model on numerical analysis and approximation theory (Section 6.1), machine learning (Section 6.2), and probability theory (Section 6.3).

6.1 Numerical Analysis and Approximation Theory

The problem considered here is to *reconstruct* a realvalued function f defined on a domain $D \subset \mathbb{R}^d$ from given *data values* $y_i = f(x_i)$ available at a set $X_n =$ $\{x_1, \ldots, x_n\}$ of distinct *data locations*. In contrast to the statistical exposition in Section 3.1, from a numerical analysis standpoint these data are not assumed to be random in any way. Nevertheless, many of the mathematical expressions that we previously motivated from a statistical perspective appear also in the solution of this numerical task. The data vector Z_n is reinterpreted as $Z_n = (f(x_1), \ldots, f(x_n))^{\top}$ and the task is to approximate the value f(x) of the unknown function f at an unsampled location $x \in D \setminus X_n$. A natural solution is a minimalnorm interpolant

$$s_{f,X_n,K} = \underset{s \in \mathcal{H}(K)}{\operatorname{arg\,min}} \|s\|_{\mathcal{H}(K)} \quad \text{s.t.} \quad \begin{array}{c} s(\boldsymbol{x}_i) = f(\boldsymbol{x}_i), \\ i = 1, \dots, n, \end{array}$$

which we recall was the third optimality property referred in Section 3. Thus, using again the *kernel matrix* $\mathbf{R}_n = [K(\mathbf{x}_i, \mathbf{x}_j)]_{i,j=1}^n$, the system $\mathbf{R}_n \mathbf{c}_n = \mathbf{Z}_n$ is solved for a fixed coefficient vector \mathbf{c}_n that determines a linear combination

$$s_{f,X_n,K}(\boldsymbol{x}) = \sum_{i=1}^n c_i K(\boldsymbol{x}_i, \boldsymbol{x}), \qquad \boldsymbol{x} \in D,$$

in the span of the *translates* $K(x_i, \cdot)$. This follows easily from the reproduction formula (3) and (2). The above formula is identical to (12) when setting $x = x_0$, and the resulting value $s_{f,X_n,K}(x)$ is interpreted as a numerical approximation to f(x). The log-likelihood function (13) can equivalently be viewed as penalising the norm of the interpolant, since $||s_{f,X_n,K}||^2_{\mathcal{H}(K)} = \mathbf{Z}_n^\top \mathbf{R}_n^{-1} \mathbf{Z}_n$.

The fourth optimality principle in Section 3 corresponds here to the fact that the norm of the error functional $\epsilon_{\boldsymbol{x}} : f \mapsto f(\boldsymbol{x}) - s_{f,X_n,K}(\boldsymbol{x})$ in the dual space $\mathcal{H}(K)^*$ of $\mathcal{H}(K)$ is minimal under all linear reconstruction algorithms in $\mathcal{H}(K)$ that use the same data \boldsymbol{Z}_n . The key tool is the *power function* P_{K,X_n} , defined for all $\boldsymbol{x} \in D$ by

$$P_{K,X_n}(\boldsymbol{x})$$

= sup { $f(\boldsymbol{x}) : f \in \mathcal{H}(K), f(X_n) = 0, ||f||_{\mathcal{H}(K)} \le 1$ }

It has the property $P_{K,X_n}(\boldsymbol{x}) = \|\boldsymbol{\epsilon}_{\boldsymbol{x}}\|_{\mathcal{H}^*(K)}$ and leads to optimal error bounds of the form

$$|f(\boldsymbol{x}) - s_{f,X_n,K}(\boldsymbol{x})| \le P_{K,X_n}(\boldsymbol{x}) ||f||_{\mathcal{H}_K}.$$

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for all $x \in D$ and $f \in \mathcal{H}(K)$. It can be numerically calculated using the kernel matrix based on $X_n \cup \{x\}$, but 685 we omit the detail. Strikingly, the power function coincides with the square root of the *kriging variance* [141], 687 giving the variance of the kriging error at x for given data 688 locations X_n and kernel K.

Analysis of the approximation error in this context thus 690 reduces to analysis of the power function, and in turn anal-991 ysis of the space $\mathcal{H}(K)$. From (4) and (23), the RKHS 692 generated by the Matérn kernel $\mathcal{M}_{\nu,1}$ has the inner prod-994 uct 694

(27)
$$\langle f,g \rangle_{\mathcal{H}(\mathcal{M}_{\nu,1})} = \int_{\mathbb{R}^d} \frac{\hat{f}(\boldsymbol{\omega})\overline{\hat{g}(\boldsymbol{\omega})}}{(1+\|\boldsymbol{\omega}\|^2)^{\nu+d/2}} \mathrm{d}\boldsymbol{\omega}$$

up to constants, which we recognise as the inner prod- 697 645 uct of the classical Sobolev space $W_2^{\nu+d/2}(\mathbb{R}^d)$. By the 698 646 Sobolev embedding theorem, the elements of this space 699 647 are well-defined continuous functions whenever $\nu > 0$. ⁷⁰⁰ 648 This space is a canonical setting for mathematical anal-701 649 ysis of PDEs, a connection that we trailed in Section 4.4. 702 650 Summarising, the use of Matérn kernels yields optimal 703 651 recovery techniques for functions in Sobolev spaces from 704 652 given sampled data. Generalised recoveries using deriva-⁷⁰⁵ 653 tive data produce *meshless* numerical methods for solving 706 654 PDEs in Sobolev spaces, including the symmetric colloca-707 655 tion method which uses derivative data for the PDE based 708 656 on Wu [172], and shares similar Hilbert space optimality ⁷⁰⁹ 657 properties Schaback [137]. The use of the Matérn kernel ⁷¹⁰ 658 is strongly motivated by the fact that PDE theory often 711 659 implies that solutions lie in Sobolev spaces. On the other ⁷¹² 660 hand, there are also good arguments to replace Matérn 713 661 714 kernels by polyharmonics [138, 50]. 662

Plenty of other results on deterministic recovery problems using kernels can be found in Wendland [166], while applications are in Schaback and Wendland [139] and MATLAB programs combined with the essential theory are in Fasshauer and McCourt [58].

In numerical analysis and approximation theory, Matérn 720 668 and other kernels are normally used for rather large val-721 669 ues of their smoothness parameter, because they seek to 722 670 solve an interpolation rather than a regression task. Nar-723 671 cowich et al. [114] proved that convergence rates then 724 672 725 depend on the minimum of the smoothness of the func-673 tion f providing the data and the kernel; a *misspecified* 674 Matérn kernel, for which the smoothness parameter ν is 675 taken to be too large relative to the smoothness of f, pro-676 duces an error that converges at the same rate as we would 677 have achieved had ν been correctly specified. On the other 678 hand, Tuo and Wang [159] prove in the same setting that 679 the prediction error becomes more sensitive to the space-680 filling property of the design points. In particular, optimal 681 convergence rates require also that the *quasi-uniformity* 682 of the experimental design is controlled. 683

Of course, the use of kernels in numerical analysis and approximation theory requires estimation of kernel parameters. The quantity σ does not arise in the correlation matrix \mathbf{R}_n , but the scale parameter α has a strong influence on the error of the interpolant. There is a vast literature on *scale estimation* that partially builds on statistical notions like ML (see references in Section 3). On the other hand, specific alternatives to the Matérn model, such as the polyharmonic kernels of Section 7.3, are able to bypass scale estimation due to the remarkable property that the interpolant is independent of the value of the scale parameter used. See Wendland [166] and Section 7.3.

6.2 Machine Learning

Kernel methods are a major strand of machine learning research, where kernels are routinely used to solve a variety of supervised and unsupervised learning tasks. Compared to the interpolatory setting of Section 6.1, data in machine learning are usually observed with noise, necessitating either a likelihood or a loss function to be specified.

The Matérn model is often convenient for the analysis of kernel methods; for example, Tuo et al. [160] provide sufficient conditions for the rates of convergence of the Matérn kernel ridge regression to exceed the standard minimax rates under both the L_2 norm and the norm of the RKHS. However, the presence of noise in the data can pose a substantial challenge to selection of smoothness parameters such as ν in the Matérn model. Karvonen [81] proves that the ML estimate of ν cannot asymptotically *undersmooth* the truth under fixed domain asymptotics; that is, if the true regression function has a Sobolev smoothness $\nu_0 + d/2$, then the smoothness parameter estimate cannot be asymptotically less than $\nu_0 + d/2$, but this in itself it not compelling motivation to use ML [82]. As a result of these additional challenges, standard practice is to keep the kernel general as far as possible when developing methodology, and as far as possible to learn a suitable form for the kernel using the data and model selection criteria. However, recent machine learning methodology for non-Euclidean data hinges on the SPDE approach, and as a consequence the Matérn and related models are explicitly being used.

As the types of data that researchers seek to analyse become more heterogeneous and structured, there has been a demand for flexible Gaussian process models defined on such non-Euclidean domains as manifolds and discrete, graph-based domains. Under the framework of Gaussian processes, Borovitskiy et al. [35] proposed to avoid numerical solution of the SPDE (24) and instead to work with a finite-rank approximation to the Gaussian process model. Specifically, they consider the SPDE in (24) appropriately adapted to a Riemannian manifold M, for which the corresponding Matérn model admits a series 774 expansion of the type 775

$$\sum_{n=0}^{\infty} \left(\frac{2\nu}{\alpha^2} + \lambda_n\right)^{-\nu - d/2} f_n(\boldsymbol{x}) f_n(\boldsymbol{x}'), \qquad \boldsymbol{x}, \boldsymbol{x}' \in M \qquad \begin{array}{c} 776 \\ 777 \\ 778 \end{array}$$

779 where $\{\lambda_n\}_{n=0}^{\infty}$ and $\{f_n\}_{n=0}^{\infty}$ are, respectively, the se-726 quences of eigenvalues and eigenfunctions from the 727 Laplace–Beltrami operator $-\Delta_M$. The authors propose 728 782 to first solve numerically for the leading eigenfunctions 729 783 ${f_i}_{i=0}^n$ of the Laplace–Beltrami operator, and then work-730 784 ing with a finite-rank Gaussian process whose realisations 731 are linear combinations of the $\{f_i\}_{i=0}^n$. Though solving 732 the eigenproblem may be harder than numerically solving 733 the SPDE, the authors argue that caching of the eigen-734 788 functions can lead to a cost saving in settings where mul-735 tiple tasks are to be solved on the same manifold. Such 736 an approach is ingeniously extended to undirected graphs 737 by Borovitskiy et al. [34], and has had a direct impact 738 792 on Gaussian processes defined on neural networks [80], 739 793 pathwise conditioning of Gaussian processes [171], sim-740 ulation intelligence in AI [93] and extension to kernel 741 methods withing graphs cross time [115]. Other applica-742 tions include Thomson sampling in neural information 796 743 systems [161], Bayesian optimisation in robotics [79], 797 744 and Gaussian processes regression on metric spaces [90]. 798 745

746 6.3 Probability Theory and Stochastic Processes

The Matérn model is well-studied from a probability 747 theory and stochastic process viewpoint. From the per-748 spective of regularity, Scheuerer [140] summarises the 749 properties of Gaussian random fields with Matérn covari-750 ance functions; sample paths are k-times differentiable in 751 the mean-square sense if and only if $\nu > k$. Under the 752 same condition, the sample paths have (local) Sobolev 753 space exponent being identically equal to k. Further, a $_{805}$ 754 Gaussian random field with Matérn covariance has frac-755 tal dimension that is identically equal to $\min(\nu, d)$, for d 756 being the dimension of the Euclidean space on which the 757 random field is defined. For non-Gaussian random fields 758 with Matérn covariance, continuity properties are studied 759 by Kent [86]. 760

Several other properties of the Matérn model have been 761 investigated. Kelbert et al. [85] study fractional random 806 762 fields under the scenario of stochastic fractional heat 807 763 equations under a Matérn model; see also Leonenko et al. 808 764 [96]. Random fields defined on the unit ball embedded 809 765 in \mathbb{R}^d , with a covariance function that is the restriction ⁸¹⁰ 766 of the Matérn model to a finite range, were studied in 811 767 Leonenko et al. [95]. Tensor-valued random fields with 812 768 an equivalent class of Matérn covariance functions were 813 769 studied in Leonenko and Malyarenko [94]. Terdik [157] 814 770 considers angular spectra for non-Gaussian random fields 771 with Matérn covariance function. A recent contribution 772 [158] provides interesting connection between the Matérn 773

model and certain Laplacian ARMA representations of a class of stochastic processes. Lilly et al. [99] show that the Matérn process is a damped version of fractional Brownian motion. Lim and Teo [100] study random fields with a generalised Matérn covariance obtained as the solution to the fractional stochastic differential equation with two fractional orders, enabling the authors to deduce the sample path properties of the associated random field. Spacetime extensions of Matérn random fields through stochastic Helmholtz equations are provided by Angulo et al. [10].

According to N. Leonenko¹, a major contributor to this literature, the importance of Matérn model is based on the Duality theorem [72, Theorem 1] which provides an explicit relation between certain classes of characteristic functions of symmetric random vectors and their density. Specifically, the spectral density associated with the Matérn model is by itself a covariance function, called the Cauchy or inverse multiquadric covariance function, that allows to parameterise the Hurst effect of the associated Gaussian random field.

This completes our tour across the scientific landscape through the lens of the Matérn model. Our attention turns now to the future, and promising enhancements that can be made to the Matérn model.

7. ENHANCEMENTS OF THE MATÉRN MODEL

This section described *enhancements* of the Matérn model; covariance functions that share (at least partially) the local properties of the Matérn model while providing additional features and functionality. Here we first introduces the models one at a time, with critical commentary on their features deferred to Section 8.

7.1 Models with Compact Support

Compactly supported covariance models have a long history that can be traced back to Askey [12], who proposed the kernel

(28)
$$\mathcal{A}_{\mu,\beta}(x) = \left(1 - \frac{x}{\beta}\right)_{+}^{\mu}, \qquad x \ge 0,$$

with β and μ being strictly positive, and where $(x)_+ = \max(0, x)$ is the *truncated power*. It was shown in that work that $\mathcal{A}_{\mu,\beta}$ belongs to Φ_d for all $\beta > 0$ if and only if $\mu \ge (d+1)/2$. Clearly, the mapping $\mathbf{x} \mapsto \mathcal{A}_{\mu,\beta}(||\mathbf{x}||)$ is compactly supported over a ball with radius β embedded in \mathbb{R}^d . As a result, covariance matrices contain exact zero entries whenever the associated states \mathbf{x}_i and \mathbf{x}_j satisfy $||\mathbf{x}_i - \mathbf{x}_j|| \ge \beta$; the computational advantages of this sparsity are discussed further in Section 8.5.

¹Personal Communication, January 2023.

Matheron's montée and descente [110] approach was 855 815 applied by Wendland [165] to the Askey functions, ob- 856 816 taining compactly supported covariance functions with 857 817 higher-order smoothness that are truncated polynomials 818 as functions of ||x||. This strategy was unable to generate 819 integer-order Sobolev spaces in even space dimensions, a 820 problem that was resolved in Schaback [136] who identi-821 fied the 'missing' Wendland functions. A unified view of 822 Wendland functions was provided by Gneiting [62]. Zas-823 tavnyi [175] provided necessary and sufficient conditions 824 for a general class encompassing both ordinary and miss-825 ing Wendland functions. Buhmann [38] provided a gener-826 alisation of Wendland functions, with sufficient paramet-827 ric conditions that allow the new class to belong to Φ_d 828 for a given d. Those functions, termed Buhmann func-829 tions, were then studied by Zastavnyi [176] and subse-830 quently by Zastavnyi and Porcu [177], Porcu et al. [129] 831 and Faouzi et al. [56]. Alternative representations and 832 properties of the Wendland functions have been studied 833 by Hubbert [76] and Chernih and Hubbert [40]. Exten-834 859 sions of the Wendland functions to multivariate [124, 46], 835 860 spatio-temporal [123] and non-stationary processes [89] 836 have also been developed. 837

A more technical discussion follows, in which we introduce two further classes of correlation functions with compact support, each of which will be the subject of discussion in Section 8.

1. The generalized Wendland (\mathcal{GW}) family [61, 176] contains correlation functions with compact support that, as in the Matérn model, admit a continuous parameterisation of smoothness of the underlying Gaussian random field. The $\mathcal{GW}_{\kappa,\mu,\beta}$ model depends on parameters $\kappa \geq 0$ and $\mu, \beta > 0$ through the identity

(29)
$$\mathcal{GW}_{\kappa,\mu,\beta}(x) = \frac{\Gamma(\kappa)\Gamma(2\kappa+\mu+1)}{\Gamma(2\kappa)\Gamma(\kappa+\mu+1)2^{\mu+1}} \mathcal{A}_{\kappa+\mu,\beta^2}(x^2) \times {}_2F_1\left(\frac{\mu}{2},\frac{\mu+1}{2};\kappa+\mu+1;\mathcal{A}_{1,\beta^2}(x^2)\right),$$

where $\mu \geq (d+1)/2 + \kappa$ is needed for $\mathcal{GW}_{\kappa,\mu,\beta}$ 842 to belong to the class Φ_d and ${}_2F_1(a, b, c, \cdot)$ is 843 the Gaussian hypergeometric function [2]. Sam-844 ple paths of the $\mathcal{GW}_{\kappa,\mu,\beta}$ model are k times mean-845 square differentiable, in any direction, if and only 846 869 if $\kappa > k - 1/2$ [61], so that κ plays the role of 847 870 the smoothness parameter in this model. When 848 $\kappa=k\in\mathbb{N},\ \mathcal{GW}_{k,\mu,eta}$ factors into the product of 849 872 the Askey function $\mathcal{A}_{\mu+k,\beta}$ with a polynomial of 850 873 degree k. This model includes the Wendland func-851 tions ($\kappa = k$, a positive integer), as well as the 852 missing Wendland functions ($\kappa = k + 1/2$). The-853 orem 1(3) in Bevilacqua et al. [25] implies that 854

2. The Gauss hypergeometric (\mathcal{GH}) family [53] is defined as

space $W_2^{\kappa}(\mathbb{R}^d)$.

(30)
$$\mathcal{GH}_{\kappa,\delta,\gamma,\beta}(x) = \frac{\Gamma(\delta - d/2)\Gamma(\gamma - d/2)}{\Gamma(\delta - \kappa + \gamma - d/2)\Gamma(\kappa - d/2)} \times \mathcal{A}_{\delta - \kappa + \gamma - d/2 + 1,\beta^2}(x^2) \times {}_2F_1\left(\delta - \kappa; \gamma - \kappa; \delta - \kappa + \gamma - d/2; \mathcal{A}_{1,\beta^2}(x^2)\right).$$

This model has four parameters and it belongs to the class Φ_d for every positive β provided $\kappa > d/2$ with

$$2(\delta - \kappa)(\gamma - \kappa) \ge \kappa$$
, and $2(\delta + \gamma) \ge 6\kappa + 1$.

Sample paths of the $\mathcal{GH}_{\kappa,\delta,\gamma,\beta}$ model are $\lceil k/2 \rceil$ times mean-square differentiable, in any direction, if and only if $\kappa > (k+d)/2$. The parameter κ thus also controls the smoothness of samples from this model.

The importance of the \mathcal{GW} and \mathcal{GH} models is discussed in Section 8.

7.2 Models with Polynomial Decay

Correlation models with polynomial decay such as the generalized Cauchy [65] or the Dagum models [22] can be useful when modelling data with long-range dependence. However, in using these correlation models one loses control over the differentiability of the the sample paths, a key property of the Matérn model. Ma and Bhadra [106] recently proposed a modification of the Matérn class that allows for polynomial decay, while maintaining the local properties of the conventional Matérn model. The correlation function associated to this model is given by

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$$\mathcal{CH}_{\nu,\eta,\beta}(x) = \frac{\Gamma(\nu+\eta)}{\Gamma(\nu)} \mathcal{U}\left(\eta, 1-\nu, \nu\left(\frac{x}{\beta}\right)^2\right), \ x \ge 0,$$

where \mathcal{U} is the confluent hypergeometric function of the second kind [2]. Here $\nu > 0$ controls mean-square differentiability near the origin, as in the Matérn case, while $\eta > 0$ controls the heaviness of the tail. The construction (31) is based on a scale mixture of (a reparameterised version of) the Matérn model involving the inverse-gamma distribution. Ma and Bhadra [106] have shown that this class is particularly useful for extrapolation problems where large distances are predominant.

75 7.3 Polyharmonic Kernels

Our catalogue of enhancements of the Matérn model finishes with *polyharmonic kernels*, defined as

(32)
$$H_{\nu,d}(x) := \begin{cases} x^{2\nu-d} \log x & \text{for } 2\nu - d \in 2\mathbb{Z} \\ x^{2\nu-d} & \text{else} \end{cases}$$

up to the sign $(-1)^{\lfloor \nu - d/2 \rfloor + 1}$. As a function of $x = \|\boldsymbol{x}\|$, 876 $x \in \mathbb{R}^d$, the Matérn kernel $\mathcal{M}_{
u-d/2,1}$ starts with even 877 powers of x followed by $H_{\nu,d}$, and in this sense the two 878 models are related. Up to a constant factor, the gener-879 alised Fourier transform of $H_{\nu,d}(||x||)$ on \mathbb{R}^d is $||\omega||^{-2\nu}$, 880 and then a scale parameter is just another constant factor. 881 This makes kernel-based interpolation by polyharmonics 882 scale-independent. Compare with (23) to see the connec-883 tion to $\mathcal{M}_{\nu-d/2,\alpha}$ in Fourier space. Stein [150] provides 884 a formal connection between polyharmonic kernels, for 885 which the name *power law* covariance functions is also 886 used, and the Matérn model. Polyharmonic kernels are 887 conditionally positive definite of order $|\nu - d/2| + 1$; 888 for a technical definition see Wendland [166]. Instead of 889 Hilbert Spaces, polyharmonic kernels generate Beppo-890 Levi spaces, which share similarities to Sobolev spaces 891 modulo that an additional polynomial space has to be 892 added to enable prediction (Section 3) and interpolation 893 (Section 6.1); see Wendland [166]. In general, polyhar-894 monic kernels arise as covariances in *fractional* Gaussian 895 fields, including forms of Brownian motion [104, Theo-896 rem 3.31. 897

Next our attention turns to a critical discussion of whether such enhancements to the Matérn model are needed.

8. ARE ENHANCEMENTS OF THE MATÉRN MODEL USEFUL?

901This final section provides critical commentary on the
921902Matérn model and the enhanced versions of the model in-
922903troduced in Section 7.

904 8.1 Rigorous Generalisation of the Matérn Model

The Matérn model does not allow for compact support, 926 905 hole effects (oscillations between positive and negative 927 906 values) at large distances, or slowly decaying tails suit-907 928 able for modeling long-range dependence. Most of the 908 enhancements in Section 7 aim to resolve these kind of is-909 sues; here we describe how the \mathcal{GW} , \mathcal{GH} and \mathcal{CH} models ₉₃₀ 910 can be viewed as rigorous generalisations of the Matérn 931 911 model. 912 932

Bevilacqua et al. [23] have shown that the Matérn $_{933}$ model is a limit case of a rescaled version of the \mathcal{GW}_{934} model. In particular they have considered the model \mathcal{GW}_{935} defined as $_{936}$

$$\widetilde{\mathcal{GW}}_{\kappa,\mu,\beta}(x) = \mathcal{GW}_{\kappa,\mu,\beta\left(\frac{\Gamma(\mu+2\kappa+1)}{\Gamma(\mu)}\right)^{\frac{1}{1+2\kappa}}}(x), \qquad x \ge 0,$$

and proved that

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$$\lim_{\mu \to \infty} \widetilde{\mathcal{GW}}_{\kappa,\mu,\beta}(x) = \mathcal{M}_{\kappa+1/2,\beta}(x), \quad \kappa \ge 0,$$

with uniform convergence over the set $x \in (0,\infty)$. The parameter μ thus allows for switching from compactly supported to globally supported models, and can either be fixed to ensure sparse correlation matrices, or can be estimated based on the dataset. However, this equivalence applies only to smoothness parameters greater than or equal to 1/2 in the Matérn model, so the full range of the smoothness parameter is not covered. This is unfortunate, since the fractal dimension [a widely used measure of roughness of the sample paths for time series and spatial data; 66] is fully parameterised using the Matérn model when the smoothness parameter lies between 0 and 1. As a consequence, the \mathcal{GW} (or \mathcal{GW}) model cannot fully parameterise the fractal dimension of the random field. This kind of issue can be solved with the \mathcal{GH} model, which includes the \mathcal{GW} model as a special case [53]:

$$\mathcal{GH}_{\frac{d+1}{2}+\nu,\frac{d+\mu+1}{2}+\nu,\frac{d+\mu}{2}+1+\nu,\beta}(x) = \mathcal{GW}_{\nu,\mu,\beta}(x)$$

Letting β , δ and γ tend to infinity in such a way that $\beta/\sqrt{4\delta\gamma}$ tends to $\alpha > 0$, the \mathcal{GH} model (30) converges uniformly to the Matérn model $\mathcal{M}_{\kappa-d/2,\alpha}(x)$, and in this case the *full range* of the smoothness parameter of the Matérn model is covered.

The Matérn model also arises as a special limit case of the CH model. Specifically, Ma and Bhadra [106] show that

$$\lim_{\eta \to \infty} \mathcal{CH}_{\nu,\eta,2\sqrt{\nu(\eta+1)\beta}}(x) = \mathcal{M}_{\nu,\beta}(x),$$

with convergence being uniform on any compact set.

The *turning band* operator of Matheron [109] can be applied to a correlation function to create hole effects while retaining positive definiteness of the kernel. An argument in Schoenberg proves that, for an isotropic correlation in \mathbb{R}^d , the correlation values cannot be smaller than -1/d [143]. Since the Matérn model is a valid model for all *d*, this implies that the application of turning bands to the Matérn model will not provide any hole effect. On the other hand, the \mathcal{GW} and \mathcal{GH} models allow for such an effect.

8.2 Estimation of Enhanced Models

ML estimation for the Matérn model are well-understood; here we discuss the extent to which similar results can be obtained for enhancements of the Matérn model.

In the context of increasing domain asymptotics, parameters of the \mathcal{GW} and \mathcal{CH} models can be estimated consistently using ML and the associated asymptotic distribution is known; see Section 3.1.1.

In the context of fixed domain asymptotics, similar to the classical Matérn model, the parameters of the these

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enhanced models cannot be consistently estimated. For instance, Bevilacqua et al. [25] show that the microergodic parameter of the covariance model $\sigma^2 \mathcal{GW}_{\kappa,\mu,\beta}$, assuming κ and μ known, is given by micro_{GW} = $\sigma^2/\beta^{2\kappa+1}$. In addition they prove that for a zero mean Gaussian field defined on a bounded infinite set $D \subset \mathbb{R}^d$ (d = 1, 2, 3), with covariance model $\sigma_0^2 \mathcal{GW}_{\kappa, \mu, \beta_0}$, the ML estimator $\hat{\sigma}_n^2/\hat{\beta}_n^{2\kappa+1}$ of the microergodic parameter is strongly consistent, *i.e.*,

$$\hat{\sigma}_n^2 / \hat{\beta}_n^{2\kappa+1} \xrightarrow{a.s.} \sigma_0^2 / \beta_0^{2\kappa+1}.$$

Additionally, for $\mu > (d+1)/2 + \kappa + 3$, its asymptotic ₉₄₅ distribution is given by 946

$$\sqrt{n}(\hat{\sigma}_n^2/\hat{\beta}_n^{2\kappa+1} - \sigma_0^2/\beta_0^{2\kappa+1}) \stackrel{d}{\longrightarrow} \mathcal{N}(0, 2(\sigma_0^2/\beta_0^{2\kappa+1})^2). \quad {}^{947}_{948}$$

949 Analogous for the \mathcal{GH} model proposed are not available 937 950 at present. 938

951 Similarly, Ma and Bhadra [106] show that the microer-952 godic parameter of the covariance model $\sigma^2 C \mathcal{H}_{\nu,\eta,\beta}$, assuming ν known, is given by

micro_{*CH*} =
$$(\sigma^2 \Gamma(\nu + \eta))/(\beta^{2\nu} \Gamma(\eta)).$$

956 In addition they prove that for a zero mean Gaussian field defined on a bounded infinite set $D \subset \mathbb{R}^d$ (d = 1, 2, 3), with covariance model $\sigma_0^2 C \mathcal{H}_{\nu,\eta_0,\beta_0}$, the ML estimator $(\hat{\sigma}_n^2/\hat{\beta}_n^{2\nu})(\Gamma(\nu+\hat{\eta}_n)/\Gamma(\hat{\eta}_n))$ of the microergodic param-960 eter is strongly consistent, *i.e.*, 961

$$\frac{\hat{\sigma}_n^2(\Gamma(\nu+\hat{\eta}_n)}{\hat{\beta}_n^{2\nu}\Gamma(\hat{\eta}_n)} \xrightarrow{a.s.} \frac{\sigma_0^2\Gamma(\nu+\eta_0)}{\beta_0^{2\nu}\Gamma(\eta_0)} -$$

and, if $\eta_0 > d/2$, its asymptotic distribution is given by

$$\frac{\hat{\sigma}_n^2(\Gamma(\nu+\hat{\eta}_n)}{\hat{\beta}_n^{2\nu}\Gamma(\hat{\eta}_n)} - \frac{\sigma_0^2\Gamma(\nu+\eta_0)}{\beta_0^{2\nu}\Gamma(\eta_0)} \\ \xrightarrow{d} \mathcal{N}\left(0, 2\left(\frac{\sigma_0^2\Gamma(\nu+\eta_0)}{\beta_0^{2\nu}\Gamma(\eta_0)}\right)^2\right).$$

These results broadly support the use of ML plug-in esti- 971 939 mates for these enhanced versions of the Matérn model; 972 940 the issue of predictive performance is discussed next. 973 941

8.3 Prediction with Enhanced Models 942

If two Gaussian measures are equivalent then the asso- 976 ciated predictions and mean squared errors are asymptot- 977 ically identical (c.f. Section 3.2). To this end, recent re- 978 sults have sought to establish equivalence between Gaus- 979 sian measures for the Matérn model and enhancements 980 of the Matérn model. Bevilacqua et al. [25] consider 981 the $\sigma_1^2 \mathcal{GW}_{\kappa,\mu,\beta}$ model and show that for given $\sigma_1 \ge 0$, 982 $\nu \geq 1/2$, and $\kappa \geq 0$, if $\nu = \kappa + 1/2$, $\mu > d + \kappa + 1/2$ and 983

(33)
$$\sigma_0^2 \alpha^{-2\nu} = \left(\frac{\Gamma(2\kappa + \mu + 1)}{\Gamma(\mu)}\right) \sigma_1^2 \beta^{-(1+2\kappa)},$$

then $P(\sigma_0^2 \mathcal{M}_{\nu,\alpha})$ is equivalent to $P(\sigma_1^2 \mathcal{GW}_{\kappa,\mu,\beta})$, for d =1, 2, 3, on the paths of $Z(\mathbf{x})$ for $\mathbf{x} \in D \subset \mathbb{R}^d$. Thus predictions made using the \mathcal{GW} model with compact support are asymptotically identical to those made using the Matérn model. Likewise, Ma and Bhadra [106] show that for a given $\eta \ge d/2$ and $\nu \ge 0$, if

(34)
$$\sigma_0^2 \alpha^{-2\nu} = \left(\frac{\Gamma(\nu+\eta)}{\Gamma(\eta)}\right) \sigma_1^2 \left(\frac{\beta^2}{2}\right)^{-\nu},$$

then $P(\sigma_0^2 \mathcal{M}_{\nu,\alpha})$ is equivalent to $P(\sigma_1^2 \mathcal{CH}_{\nu,\eta,\beta})$, for d =1,2,3, on the paths of Z(x) for $x \in D \subset \mathbb{R}^d$. Thus predictions made using the \mathcal{GW} model with polynomial tail decay are asymptotically identical to those made using the Matérn model.

If interest is in the predictor (12), but not the predictive uncertainty resulting from the associated Gaussian random field, then it is interesting to note that the stationarity assumption of the Matérn model may not be needed. Stein et al. [153] showed that, under suitable parametric conditions, one can consider $\alpha = 0$ in the Matérn model, and this is equivalent to prediction using the polyharmonic kernels $H_{\nu,d}$ in (32). Theorem 1 in that work shows that if $d \leq 3$ and the parameter ν satisfies condition (2) therein (or d = 1), then it is impossible to distinguish $\alpha > 0$ from $\alpha = 0$ on a bounded domain. The above observation reflects the fact that prediction using polyharmonic kernels, like in Section 6.1, is scale-independent. This follows from homogeneity of the Fourier transform and eliminates the need for scale estimation in this context.

8.4 Screening with Enhanced Models

The screening effect extends also to enhanced versions of the Matérn model. For regular schemes, Theorem 1 in Porcu et al. [128] shows that the \mathcal{GW} model allows for an asymptotic screening effect when $\mu > (d+1)/2 + \kappa$. This condition is not restrictive, since $\mu \ge (d+1)/2 + \nu$ is already required for $\mathcal{GW}_{\kappa,\mu,\beta}$ to belong to the class Φ_d . For irregular schemes the situations is more complicated. For example, for non-differentiable fields in d = 1, Theorem 1 in Stein [151] in concert with Theorem 1 in Porcu et al. [128] explains that the Askey model $\mathcal{GW}_{0,\mu,\beta}$ allows for a screening effect provided $\mu > 1$. For d = 2, Theorem 2 in Stein [151] implies that the Askey model allows for screening provided that $\mu > 3/2$. The \mathcal{GW} model satisfies Stein's condition in (1.3) of Porcu et al. [128], which in turn allows the Stein hypothesis (22) to be verified.

The numerical experiments in Porcu et al. [128] suggest that the screening effect is even stronger under enhanced models with compact support, compared to the standard Matérn model. This can deliver computational advantages, which we discuss next.

985 8.5 Scalable Computation

1038 The eternal fight between statistical accuracy and com-986 1039 putational scalability has produced methods that attempt 987 to deal with this notorious trade-off. The discussion that 988 041 follows focuses specifically on this trade-off in the con-989 1042 text of the Matérn model. General approaches, such as 990 1043 those based on predictive processes [17] and those based 991 on fixed-rank kriging [45], will not be discussed; the in-992 1045 terested reader is referred to the review of Sun et al. [154]. 993 1046

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The computational complexity associated with the Matérn model is broadly governed by the input space dimension (d), the number (p) of kernel parameters that must be estimated, and the number of data (n). These challenges will be considered in turn.

First we consider the challenge of large d, which is of-999 1052 ten encountered in machine learning, when Gaussian pro-1000 cess regression is performed on high-dimensional input 1001 spaces [169]. Since the Matérn model $\mathcal{M}_{\nu,\alpha}$ reproduces a 1055 1002 Sobolev space up to an equivalent norm (c.f. Section 6.1), $\frac{1}{1056}$ 1003 and $\nu > d/2$ is required to for elements of this space to 1057 1004 be pointwise well-defined, it follows that ν must tend to 1058 1005 infinity as d tends to infinity, so that the Matérn model re-1006 duces to the Gaussian model (10). The flexibility of some $_{1060}$ 1007 enhanced models is also lost in this limit; the condition 1061 1008 $\mu \geq (d+1)/2 + \kappa$ in the $\widetilde{\mathcal{GW}}_{\kappa,\mu,\beta}$ model forces the pa-1009 rameter μ to go to infinity with d, which in turn forces 1063 1010 $\mathcal{GW}_{\kappa,\mu,\beta}$ to approach $\mathcal{M}_{\nu,\alpha}$. From this point of view the 1064 1011 class $\mathcal{GH}_{\kappa,\delta,\gamma,\beta}$ seems more promising to use for large 1065 1012 d. An additional remark is that, for d > 5, all Gaussian 1013 measures with Matérn covariance functions are orthogo-1014 nal [8]. This has philosophical consequences for Gaussian 1015 process regression when the Matérn model is viewed as a 1016 prior distribution encoding a priori belief, since a small 1017 change to the kernel parameters results in the entire sup-1018 port of the prior being changed. 1019

Coupled to large input dimension d is the challenge 1020 where there are a large number of parameters p appearing 1021 in the model. The multivariate Matérn model suffers from 1066 1022 the fact, not only does p increase exponentially with d, ¹⁰⁶⁷ 1023 but the conditions for validity of the model imply severe 1068 1024 restrictions on the collocated correlation coefficient ρ_{ii} in ¹⁰⁶⁹ 1025 (26). Emery et al. [54] show that such restrictions become ¹⁰⁷⁰ 1026 extremely severe already with p = 3. Similar comments ¹⁰⁷¹ 1027 apply to other multivariate covariance functions, includ-1072 1028 ing the multivariate \mathcal{GW} model in Daley et al. [46]. 1029

Finally we consider the case where the number n of 1074 1030 1075 data is large, entailing a $O(n^3)$ computational and $O(n^2)$ 1031 1076 storage cost associated with the predictor (12). Several ap-1032 proaches have been proposed to reduce these costs in the ¹⁰⁷⁷ 1033 context of the Matérn model, many of which take advan-1078 1034 tage of the (approximate) sparsity of the covariance $(\Sigma_n)^{1079}$ 1035 or precision (Σ_n^{-1}) , or its Cholesky factor $(ch(\Sigma_n^{-1}))$: 1080 1036 1081

- Sparsity in the covariance matrix Σ_n of the Matérn model is directly exploited by enhanced versions of the Matérn model from Section 7. Such approaches can be useful when the (estimated) compact support is relatively small with respect to the spatial extent of the sampling region, so that approximations are extremely sparse; see below for an empirical investigation of this point.
- The precision matrix Σ_n⁻¹ associated with the Matérn model is in general non-sparse (except for the case d = 1 and ν = 0.5) but it turns out that the matrix values are in general relatively close to 0, i.e. Σ_n⁻¹ is quasi-sparse. As a consequence, approximating Σ_n⁻¹ with a sparse matrix can be a good strategy. A notable instance of this approach is the SPDE approach from Section 4.2. This approach can be also motivated from results in numerical linear algebra, which demonstrate that if the elements of a matrix show a property of decay, then the elements of its inverse also show a similar (and faster) behavior [20].
- Vecchia's approximation [163] and its extensions [e.g. 48, 67, 83, 47] imply a sparse approximation of of ch(Σ_n⁻¹) and are often applied to the Matérn model, although they can be applied to any covariance model. One potential limitation of these method is that they depend on an ordering of the variables and the choice of conditioning sets which determines the Cholesky sparsity pattern [see 67].

It is instructive to numerically investigate the sparseness of matrices associated with enhancements of the Matérn model, and for this we focus on the $\widetilde{\mathcal{GW}}_{\kappa,\mu,\beta}$ model, which allows us to switch from a model with compact support of radius

$$C = \beta \left(\frac{\Gamma(\mu + 2\kappa + 1)}{\Gamma(\mu)} \right)^{\frac{1}{1 + 2\kappa}}$$

to the Matérn model by increasing the μ parameter. In our experiment, the sparseness of Σ_n and the *quasisparseness* of Σ_n^{-1} and $ch(\Sigma_n^{-1})$ are reported, the latter being defined as the percentage of values in the upper triangular matrix with absolute value lower than an arbitrary small constant ϵ , and in our example we set $\epsilon = 1.e - 8$.

The empirical assessment considers n = 1,156 and n = 4,489 location sites over $[0,1]^2$, where the points are equally spaced by 0.03 and 0.015 respectively in a regular grid. For $\nu = 0,1,2$, we set β such that the practical range of the Matérn model is equal to 0.15 ($\beta = 0.050, 0.0316, 0.0253$ respectively), and consider increasing $\mu = 1.5 + \kappa, 4, 8, 16, 32, 120, \infty$ (with $\widetilde{\mathcal{GW}}_{\kappa,\infty,\beta}$ being the Matérn model $\mathcal{M}_{\kappa+1/2,\beta}$).

The results are reported in Table 1. For the low values $\mu = 1.5, 2.5, 3.5$ and $\nu = 0, 1, 2$, the covariance matrix is highly sparse, while the sparseness decreases when

increasing μ , as expected. There is a clear trade-off be-tween the sparseness of Σ_n and quasi-sparseness of Σ_n^{-1} and $ch(\Sigma_n^{-1})$ for each $\nu = 0, 1, 2$. However, when increas-ing μ , that is when Σ_n approaches the Matérn covariance

As in Table 1, but with β chosen such that the practical range of the Matérn model is equal to 0.4.

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		$\kappa = 0$										κ	= 1							κ		I		
	μ C Σ_n			Σ	$\frac{-1}{n}$	$\mathrm{ch}(\boldsymbol{\Sigma}_n^{-1})$			C	Σ_n		$\mathbf{\Sigma}_n^{-1}$		$\mathrm{ch}(\mathbf{\Sigma}_n^{-1}) \qquad \mu$		μ	C	$\mathbf{\Sigma}_n$		$\mathbf{\Sigma}_n^{-1}$		$\operatorname{ch}(\mathbf{\Sigma}_n^{-1})$		
			1156	4489	1156	4489	1156	4489			1156	4489	1156	4489	1156	4489			1156	4489	1156	4489	1156	4489
ī	1.5	0.20	90.0	89.1	0	0	0	0	2.5	0.28	80.5	80.2	0	0	0	0	3.5	0.35	72.6	71.4	0	0	0	0
	4	0.53	48,7	47.3	1.03	9.62	0.58	2.77	4	0.42	65.0	63.7	0	1.03	1.12	0.81	4	0.39	67.2	66.4	0	0	0	0
	8	1.07	1.43	1.21	4.14	25.6	1.84	8.15	8	0.75	20.7	19.7	5.91	14.5	4.49	12.9	8	0.67	30.7	29.6	0.71	2.73	2.13	6.90
	16	2.13	0	0	15.0	43.5	4.98	21.3	16	1.43	0	0	21.2	45.2	17.1	30.8	16	1.21	1.72	1.29	13.7	25.8	14.8	27.9
	32	4.27	0	0	22.8	44.4	12.7	24.3	32	2.78	0	0	37.2	55.0	29.0	41.0	32	2.29	0	0	25.5	23.5	24.8	42.7
	120	16.0	0	0	21.0	46.2	8.33	21.9	120	10.2	0	0	40.8	59.8	32.2	46.2	120	8.24	0	0	26.0	10.9	25.9	46.4
	∞	∞	0	0	23.4	47.9	9.85	24.0	∞	∞	0	0	42.6	61.3	34.3	48.0	∞	∞	0	0	26.0	18.2	25.9	46.6

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range of the Matérn model is equal to 0.15.

the $\widehat{\mathcal{GW}}_{\kappa,\mu,\beta}$ model. The case $\widehat{\mathcal{GW}}_{\kappa,\infty,\beta}$ corresponds to the Matérn 1083 model $\mathcal{M}_{\nu+1/2,\beta}$. The β parameters are chosen so that the practical 1084 Sparsity (percentage of zero values in the upper triangular part) of the covariance matrix Σ_n , and quasi-sparsity (defined in the main text) in the precision matrix (Σ_n^{-1}) and its Cholesky factor $(ch(\Sigma_n^{-1}))$ for TABLE 1 1084

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I	$\kappa = 0$								$\kappa = 1$								$\kappa = 2$							
	μ	$\mu C \Sigma_n \Sigma_n^{-1} \operatorname{ch}(\Sigma_n^{-1}) \mu$				μ	C Σ_n Σ_n^{-1}						$(2n^{-1})$	μ	C	Σ_n Σ_n^{-1}			$\frac{-1}{2}$	1 ch (Σ_n^{-1})				
Ī		1156 4489 1156 4489 1156 4489		4489		1156 4489			1156 4489 1156 4489			4489			1156 4489 1		1156	156 4489 115		4489				
ī	15	0.07	08.4	08.3	30.3	1.46	35.7	9.17	25	0.11	07.1	06.7	3.80	1.67	6 30	4 44	35	0.13	04.7	95.0	2.00	1 91	5.63	6 1 2
	4	0.20	90.4 90.1	89.6	45.9	56.0	45.0	54.9	4	0.11	93.4	93.3	34.2	47.1	36.6	48.7	4	0.15	94.7 94.8	94.2	10.4	11.9	15.0	22.6
	8	0.40	66.4	65.7	56.6	64.2	52.0	60.0	8	0.28	80.9	80.7	69.7	57.0	69.3	59.2	8	0.25	84.7	84.4	42.8	61.0	47.3	64.8
	16	0.80	16.4	15.2	59.8	71.5	53.8	66.6	16	0.54	48.1	47.1	64.2	80.1	64.0	76.0	16	0.45	59.1	58.3	51.7	73.0	54.6	74.2
	32	1.60	0	0	61.2	79.0	55.8	72.8	32	1.04	1.91	1.61	66.4	83.8	66.5	82.2	32	0.86	11.0	10.1	52.3	75.7	54.8	77.2
	120	6.01	0	0	65.2	76.6	58.9	68.2	120	3.82	0	0	66.2	84.6	65.9	81.2	120	3.09	0	0	52.4	75.7	54.1	76.5
	∞	∞	0	0	66.2	77.5	58.9	70.0	∞	∞	0	0	66.2	84.9	65.1	80.7	∞	∞	0	0	52.4	75.4	53.6	74.6

matrix, then Σ_n^{-1} or ch (Σ_n^{-1}) tends to be highly quasi-1137 sparse.

We replicate the same experiment but with a practical ¹¹³⁹ range of the Matérn model equal to 0.4. This leads to ¹¹⁴⁰ $\beta = 0.133, 0.084, 0.067$ for $\nu = 0, 1, 2$ respectively. The ¹¹⁴¹ results are reported in Table 2. The conclusions are the ¹¹⁴³ same of the previous setting but in this case, we have ¹¹⁴⁴ lower levels of sparseness for Σ_n and of quasi-sparseness ¹¹⁴⁵ for Σ_n^{-1} and ch(Σ_n^{-1} .

These numerical experiments highlight a clear trade-off 1147 1096 between the (quasi-)sparseness of Σ_n^{-1} (or ch($\Sigma_n^{-1})$) and $_{_{1149}}^{_{_{1149}}}$ 1097 Σ_n when increasing μ for fixed β and ν i.e. when switch-1150 1098 ing from a compactly supported to a globally supported 1151 1099 Matérn model. In particular, when $\mu \to \infty$ (the Matérn¹¹⁵² 1100 model), then Σ_n^{-1} is highly quasi-sparse and Σ_n is dense. ¹¹⁵³ In contrast, when μ is small then Σ_n^{-1} is not quasi-sparse ¹¹⁵⁴₁₁₅₅ 1101 1102 yet Σ_n is highly sparse. This seems to suggest that sparse $_{1156}$ 1103 precision matrix approximation should work reasonably 1157 1104 well for the Matérn model, but could be problematic when 1158 1105 handling data exhibiting short compactly supported de-1159 1106 pendence. In this case a better approach should be to ex-1107 ploit the sparsity of Σ_n , as enabled by enhanced versions $\frac{1}{162}$ 1108 of the Matérn model. 1109 1163

9. CONCLUSION

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The impact of the Matérn model since its conception 1167 1110 has been substantial, and the model continues to be widely 1168 1111 used, across a broad range of scientific disciplines and be-1169 1112 yond. While the original motivation for the Matérn model ¹¹⁷⁰ 1113 came from its flexibility in context of spatial interpola- $\frac{1171}{1172}$ 1114 tion, there is now also a rich literature of alternative and 1173 1115 enhanced versions of the model. In particular, the SPDE 1174 1116 and related approaches enable one to define analogues of 1175 1117 the Matérn model on quite general domains, admitting 1176 1118 sparse approximations to precision matrices, while recent $\frac{1177}{1178}$ 1119 advances in enhanced models with compact support can $\frac{1}{1179}$ 1120 facilitate scalable computation through sparse approxima-1180 1121 tion of covariance matrices, and are well-suited to pro-1181 1122 cesses with short-scale dependence. The theoretical and 1182 1123 empirical properties of these enhanced models have been 1183 1124 recently and actively studied. On the other hand, there re-1125 main open theoretical issues of practical importance, such 1186 1126 as parameter estimation at finite sample sizes, and the im-1187 1127 pact of parameter estimation on the performance of the 1188 1128 associated predictions. 1189 1129

Our current understanding of the Matérn model has ¹¹⁹⁰ emerged as the result of engagement between scientists ¹¹⁹¹ and practitioners from different disciplines, and our hope ¹¹⁹³ is that this multi-disciplinarity perspective will shine fur- ¹¹⁹⁴ ther light onto the Matérn model. ¹¹⁹⁵

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MATÉRN: A JOURNEY

Supplementary Material

APPENDIX A: MODELLING THROUGH MATÉRN IN UNCONVENTIONAL SCENARIOS: THE EXTENDED VERSION

One might object that the Matérn class is limited to scalar-valued random fields that are stationary and isotropic. While this being true, it is also true that the Matérn class represents the building block for way more sophisticated scenarios. We list them throughout.

- 1691 1. Scalar-valued random fields.
- 1692a) Anisotropies. If spatial dimension develops over preferential directions, isotropy is no longer a realistic1693assumption for spatial modeling. Several types of anisotropies are challenged in [7], and it is shown that the1694Matérn class can be composed with *ad hoc* deformations so to take into account preferential directions in1695terms of spatial dependence.
 - b) Nonstationarity. The Matérn kernel has been used by [119] to build nonstationary models. Consider a collection of Gaussian distributions indexed by their mean, such that the element with mean x has covariance matrix Σ_x . Let

$$Q_{\boldsymbol{x},\boldsymbol{y}} = (\boldsymbol{x} - \boldsymbol{y})^{\top} \left(\frac{\Sigma_{\boldsymbol{x}} + \Sigma_{\boldsymbol{y}}}{2}\right)^{-1} (\boldsymbol{x} - \boldsymbol{y})$$

and

$$K(\boldsymbol{x},\boldsymbol{y}) = \left| \Sigma_{\boldsymbol{x}} \right|^{1/4} \left| \Sigma_{\boldsymbol{y}} \right|^{1/4} \left| \frac{\Sigma_{\boldsymbol{x}} + \Sigma_{\boldsymbol{y}}}{2} \right|^{-1/2} \mathcal{M}_{\nu,\alpha} \left(\sqrt{Q_{\boldsymbol{x},\boldsymbol{y}}} \right).$$

Then, K is positive definite. This approach was recently generalised in Roininen et al. [132]. Alternatively, one may induce non-stationarity by *warping* the inputs of the Matérn covariance function, as $K(x, y) = \mathcal{M}_{\nu,\alpha}(||w(x) - w(y)||)$ for some diffeomorphism $w : \mathbb{R}^d \to \mathbb{R}^d$ that may itself be parametrised. The case in which w is parametrised by a deep neural network was explored in Wilson et al. [170].

- c) Graphs and Quasi Metric spaces. [9] consider graphs with Euclidean edges, equipped with either the geodesic distance over the graph, or the resistance metric. They prove that $\mathcal{M}_{\nu,\alpha}$ can be composed with the resistance metric over the graph provided $0 < \nu \leq 1/2$. More recently, [112] provide a generalisation of this setting by considering quasi-metric spaces. Apparently, similar restrictions hold for this case. Recently, [32] adopt a different approach to build random fields with their covariance structure on metric graphs. Space-time version of the Matérn class, for space being a graph with Euclidean edges, have been considered by [155] and by [127].
 - d) Space-time. For a space-time Gaussian random field $\{Z(\boldsymbol{x},t), \boldsymbol{x} \in \mathbb{R}^{d+1}, t \in \mathbb{R}\}$, a typical second-order assumption is that the covariance is isotropic in space and stationary over time. That is,

(35)
$$\operatorname{Cov}\left(Z(\boldsymbol{x},t),Z(\boldsymbol{y},t')\right) = K\left(\|\boldsymbol{x}-\boldsymbol{y}\|,|t-t'|\right)$$

for all $(\boldsymbol{x}, t), (\boldsymbol{y}, t') \in \mathbb{R}^d \times \mathbb{R}$. The Matérn function has been used as a building block for such a structure. Of particular interest are *non-separable* covariance functions, that allow for an interaction between space and time, and in this context [62] and [178] prove that

(36)
$$K(x,u) = \frac{\sigma^2}{\psi(u^2)} \mathcal{M}_{\nu,\alpha}\left(\frac{x}{\psi(u^2)}\right), \ x, u \ge 0,$$

generate a valid space-time covariance function of the type (35). Here, ψ is a strictly positive function having a completely monotonic derivative [21].

e) Non Gaussian Fields. Through consideration of transformations $w: \mathbb{R} \to \mathbb{R}$, one can use the Matérn model 1710 as the basis for a range of non-Gaussian models $Z(\mathbf{x}) = w(Z(\mathbf{x}))$. However, the covariance function of Z 1711 will not be Matérn in general. The question of whether there exist non-Gaussian processes whose covariance 1712 function is nevertheless of Matérn class was answered positively in [1]. [174] have proposed trans-Gaussian 1713 random fields with Matérn covariance function. [29] and subsequently [164] have provided SPDE based 1714 constructions for non Gaussian Matérn fields. A general class of non Gaussian fields with kernel $g(\mathcal{M}_{\nu,\alpha})$ 1715 with $g(\cdot)$ a suitable function that preserves positive definiteness can be obtained through a transformation of 1716 (independent replicates of) a Gaussian field (see for instance [120, 173, 24, 113]). 1717

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- f) Mixtures of Gaussian Fields. One can model a real-world phenomenon as a superposition Z₁ + Z₂ where Z₁ might be selected to capture an overall trend (e.g. Z₁(x) = c₁φ₁(x) + ··· + c_pφ_p(x) for some fixed functions φ₁,...,φ_p and some coefficients c ∈ ℝ^p and Z₂ could be Matérn, to capture the level of smoothness of the process. Such processes appear in Bayesian linear regression, where the coefficient c is viewed as random due to epistemic uncertainty, for example c ~ N(μ, Σ), and both c and Z₂ are to be jointly inferred. The resulting mixture Z₁ + Z₂ is then again a Gaussian random field.
 - g) Matérn Models with Modified Tails. The Matérn covariance function decays exponentially with distance. This can be a drawback in the presence of long memory. Some classes of covariance functions allow to index long versus short memory in spatial data. The generalised Cauchy [65] and the Dagum [22] are prominent examples of families allowing to index both fractal dimensions and long memory, also termed Hurst effect. Yet, these families do not allow to parameterise the smoothness in the same fashion of the Matérn class. This dilemma has preoccupied several scientists. Below we describe the approaches devoted to modify the tails of the Matérn model while preserving (a) positive definiteness and (b) local behavior, in turn connected with mean square differentiability and Sobolev space parameterisation. Each of the contributions below has different motivations as explained throughout.
 - (a) Matérn models with periodic tails. A modification of the spectral density of the Matérn model has been proposed by [91]. The primary idea is to propose an isotropic class of spectral densities, $\widetilde{M}_{\nu,\alpha,\sigma^2}$, that is connected to the Matérn family $\widehat{M}_{\nu,\alpha,\sigma^2}$ through the identity

(37)
$$\widetilde{M}_{\nu,\alpha,\sigma^2}(z) = \left(b^2 + z^2\right)^{\xi} \widehat{\mathcal{M}}_{\nu,\alpha,\sigma^2}(z),$$

with $b \ge 0$ and $\xi < \nu$. While b is an additional range parameter, ξ is related to the smoothness of the respective process. More precisely, the random field is k-times mean square differentiable if and only if $\nu - \xi > k$. In the limit case $\xi \to 0$, the traditional Matérn model is recovered.

This scale in the spectral density produces a shift in the mode of the spectrum; thus, it is particularly useful to obtain processes with strong periodicities. The covariance function associated to (37) does not have a known explicit expression, so statistical methodologies in the spectral domain should be employed when dealing with this model.

(b) *Hybrid Models*. Another generalisation of the Matérn model exploits the fact that it can be written as a scale mixture of a Gaussian kernel against a probability density function

(38)
$$\mathcal{M}_{\nu,\alpha}(x) = \int_0^\infty \exp(-ux^2)\pi_{IG}(u;\nu,1/(4\alpha^2))\mathrm{d}u$$

of the inverse gamma type. [5] make use of the identity above to create *hybrid* models that allow to preserve the local properties of the conventional Matérn model while attaining more flexible behaviours at large distances. One potential hybrid construction, called Matérn-Cauchy and denoted $\mathcal{MC}_{\nu_1,\nu_2,\alpha,\xi}$, is attained through

(39)
$$\mathcal{MC}_{\nu_1,\nu_2,\alpha,\xi}(x) = \int_0^\xi \exp(-ux^2)\pi_G(u;\nu_1/2,\alpha)\mathrm{d}u + \int_\xi^\infty \exp(-ux^2)\pi_{IG}(u;\nu_2,1/(4\alpha^2))\mathrm{d}u.$$

Here, π_G is the gamma probability density function, with a shape-rate parameterisation, π_{IG} is the density in (41), $\alpha > 0$ is a range parameter, $\nu_1 > 0$ controls the polynomial rate of decay of the covariance, $\nu_2 > 0$ indexes the mean square differentiability, and $\xi \ge 0$ is an additional parameter that balances the Matérn and Cauchy contributions to the total covariance function. As $\xi \to 0$, the hybrid model tends to a Matérn covariance. A closed form expression for this model is provided in [5]. Note that this covariance function is positive definite in any dimension, and is a natural competitor of the model (31).

Another hybrid model is constructed in [5] by replacing the Gaussian kernel in (38) with a difference of Gaussian kernels

$$(a \exp(-ubx^2) - \exp(-ux^2))(a-1)^{-1}$$

where *a* and *b* satisfy the condition $1 < b < a^{2/d}$, in order to obtain a positive definite kernel in \mathbb{R}^d [130]. The resulting model has a local behavior of Matérn type and allows for negative correlations at large distances (hole effect). The parameters *a* and *b* control the sharpness of the hole effect. When *a* is arbitrarily large, the conventional Matérn model is recovered. Algebraically closed expressions are reported in [5].

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2. Vector-valued random fields.

Let $\{Z(x), x \in \mathbb{R}^d\} \subset \mathbb{R}^p$ be a *p*-variate random field with isotropic covariance mapping $K : \mathbb{R}^d \to \mathbb{R}^{p \times p}$ having elements K_{ij} defined as

$$K_{ij}(\boldsymbol{x}) = \operatorname{cov}\left(Z_i(\boldsymbol{0}), Z_j(\boldsymbol{x})\right), \, \boldsymbol{0}, \boldsymbol{x} \in \mathbb{R}^d.$$

Specifically, K_{ii} and K_{ij} are termed auto and cross-covariance function.

a) *Multivariate Spatial*. There has been a plethora of approaches related to multivariate spatial modeling, and the reader is referred to [60]. [64] have proposed a multivariate covariance structure of the type

$$K_{ij}(\boldsymbol{x}) = \sigma_{ii}\sigma_{jj}\rho_{ij}\mathcal{M}_{\nu_{ij},\alpha_{ij}}(\|\boldsymbol{x}\|), \, \boldsymbol{x} \in \mathbb{R}^{d},$$

where σ_{ii}^2 is the variance of Z_i from Z, and ρ_{ij} is the collocated correlation coefficient. There are restrictions on the parameters ν_{ij} , α_{ij} and ρ_{ij} to preserve positive definiteness, and often the restrictions on the collocated correlations coefficients ρ_{ij} are severe. This motivated alternative approaches to alleviate the parametric restrictions, and the reader is referred to [11] and more recently to [54].

b) *Multivariate space-time*. The setting above can be generalised to space-time by considering $\{Z(x,t), x \in \mathbb{R}^d, t \in \mathbb{R}\} \subset \mathbb{R}^p$ a *p*-variate random field with isotropic covariance mapping $K : \mathbb{R}^d \times \mathbb{R} \to \mathbb{R}^{p \times p}$ having elements K_{ij} defined as

$$K_{ij}(\boldsymbol{x}, u) = \operatorname{cov}\left(Z_i(\boldsymbol{0}, t), Z_j(\boldsymbol{x}, t+u)\right)$$

for $\mathbf{0}, \mathbf{x} \in \mathbb{R}^d$, $t, u \in \mathbb{R}$. [36] consider mappings K_{ij} of the type

$$K_{ij}(\boldsymbol{x}, u) = \frac{\sigma_{ii}\sigma_{jj}\rho_{ij}}{\psi(u^2)^{d/2}}\mathcal{M}_{\nu_{ij}, \alpha_{ij}\psi(u^2)}(\|\boldsymbol{x}\|)$$

for $x \in \mathbb{R}^d$, $u \in \mathbb{R}$ and a suitable positive valued and continuous function ψ . This setting has been recently generalised by [6] and through a technical approach by [125]: for both contribution, the idea is to replace (pointwise) the mapping ψ with the mapping ψ having continuous and strictly positive elements ψ_{ij} .

c) Multivariate Nonstationary. [88] derive a class of matrix valued covariance functions where the direct and cross-covariance functions belong to the Matérn class. The parameters of the Matérn class are allowed to vary with location, yielding local variances, local ranges, local geometric anisotropies and local smoothnesses. Define Σ_{i,x} = Σ_i(x) : ℝ^d → ℝ^{d×d} and assume Σ_i(x) is positive definite for all i and all x ∈ ℝ^d. Let

$$Q_{ij;\boldsymbol{x},\boldsymbol{y}} = (\boldsymbol{x} - \boldsymbol{y})^{\top} \left(\frac{\Sigma_{i,\boldsymbol{x}} + \Sigma_{j,\boldsymbol{y}}}{2}\right)^{-1} (\boldsymbol{x} - \boldsymbol{y})$$

and

$$K_{ij}(\boldsymbol{x},\boldsymbol{y}) = \rho_{ij}\sigma_{i,\boldsymbol{x}}\sigma_{j,\boldsymbol{y}}\mathcal{M}_{\nu,\alpha}\left(\sqrt{Q_{ij;\boldsymbol{x},\boldsymbol{y}}}\right)$$

Then, $\boldsymbol{K}(\cdot, \cdot) = [K_{ij}(\cdot, \cdot)]_{i,j=1}^p$ is positive definite.

d) Multivariate Matérn with Dimple. In a bivariate spatial context, each element of the matrix-valued Matérn covariance function admits a scale-mixture representation as in (38). [4] considered a modification of such a mixture to obtain a generalization of the Matérn model, given by

(40)
$$\widetilde{K}_{ij}(x) = \sigma_{ii}\sigma_{jj}\rho_{ij}\int_0^\infty \exp(-ux^2)g_{ij}(u;\xi) \times \pi_{IG}(u;\nu_{ij},1/(4\alpha_{ij}^2))\mathrm{d}u,$$

where $g_{ii}(u;\xi) = 1$, and $g_{ij}(u;\xi) = 1\{u \le \xi\} - 1\{u \ge \xi\}$ for $i \ne j$, with $1\{\cdot\}$ being the indicator function and ξ a nonnegative parameter, and where π_{IG} is the probability density function of an inverse gamma random variable, that is

(41)
$$\pi_{IG}(z;a,b) = \frac{b^a}{\Gamma(a)} z^{-a-1} \exp(-b/z), \ z > 0.$$

Observe that the diagonal elements of the matrix-valued covariance are not altered; thus, the appealing local attributes of the Matérn model are maintained. This construction only has an impact on the cross-covariances. Indeed, [4] showed that $\widetilde{K}_{12}(x)$ is not a monotonically decreasing function of x. More precisely, the cross-covariance can attain its maximum value at a strictly positive distance. This property was called cross-dimple in [4]. The parameter ξ regulates the intensity of the cross-dimple. Clearly, the traditional bivariate Matérn model is a limit case of this construction ($\xi \to \infty$). Closed-form expressions for (40) are provided by [4]. Moreover, for $\nu_{12} = 1/2 + n$, $n \in \mathbb{N}$, $\widetilde{K}_{12}(x)$ can be expressed in terms of error functions and exponential functions.

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- Birections, shapes and curves. The Matérn model is central to the study of directional processes. [18] formalize the notions of directional finite difference processes and directional derivative processes with special emphasis on the Matérn covariance function. They provide complete distribution theory results under the assumptions of a stationary Gaussian process (with Matérn covariance) model either for the data or for spatial random effects.
- [16] introduced Bayesian wombling to measure gradients related to curves through wombling boundaries. The
 smoothness properties of the Matérn model are proved to be successful within such a framework. Modeling approaches to temporal gradients using the Matérn model have been proposed by [131].